



# Nonlocal transport in the scrape-off layer

**P. Manz<sup>1,2</sup>, F. Fischer<sup>1</sup>, G. Birkenmeier<sup>1,2</sup>, D. Carralero<sup>2</sup>,  
G.Fuchert<sup>2</sup>, S. Marsen<sup>3</sup>, B. Nold<sup>4</sup>, M. Ramisch<sup>4</sup>,  
T.T. Ribeiro<sup>2</sup>, B.D.Scott<sup>2</sup>, U.Stroth<sup>2,1</sup>, R.S. Wilcox<sup>5</sup>**

<sup>1</sup>Technische Universität München, Germany

<sup>2</sup>Max-Planck-Institut für Plasmaphysik, Garching, Germany

<sup>3</sup>Max-Planck-Institut für Plasmaphysik, Greifswald, Germany

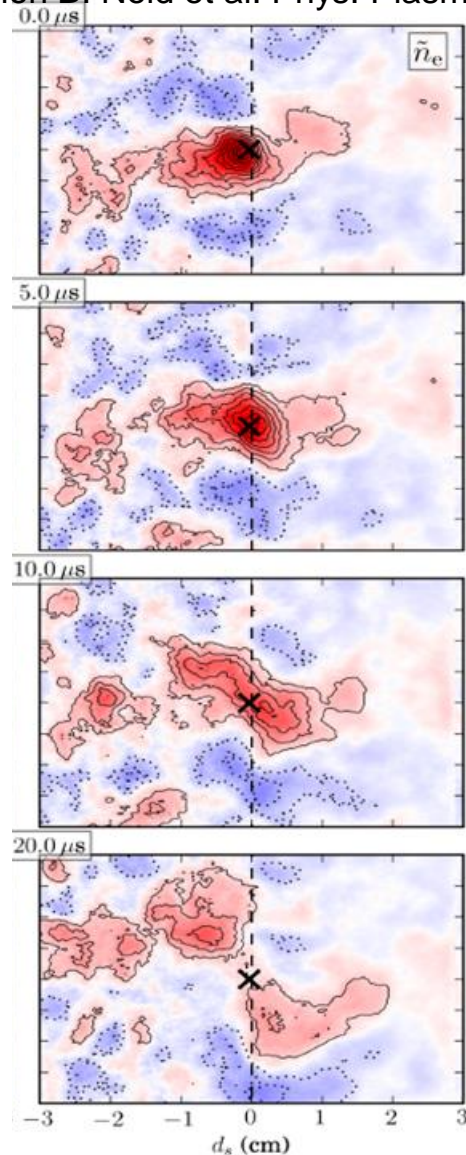
<sup>4</sup>Institut für Grenzflächenverfahrenstechnik und Plasmatechnologie,  
Universität Stuttgart, Germany

<sup>5</sup>Oak Ridge National Laboratory, OakRidge TN, USA

- Conditional averaged picture at the shear layer
- Statistics of blob trajectories
- Relation between turbulence spreading and blob generation
- Results from HSX, TJ-K, WEGA

# Blob generation

GEMR simulation B. Nold et al. Phys. Plasmas 2014



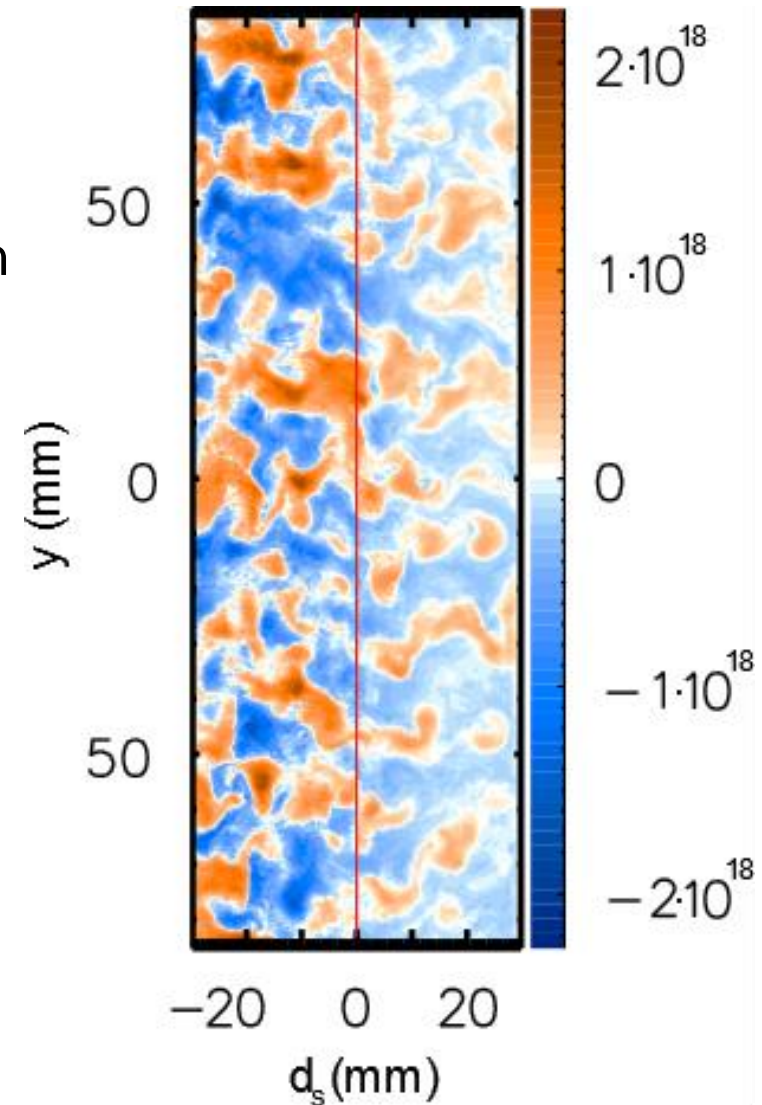
theory:

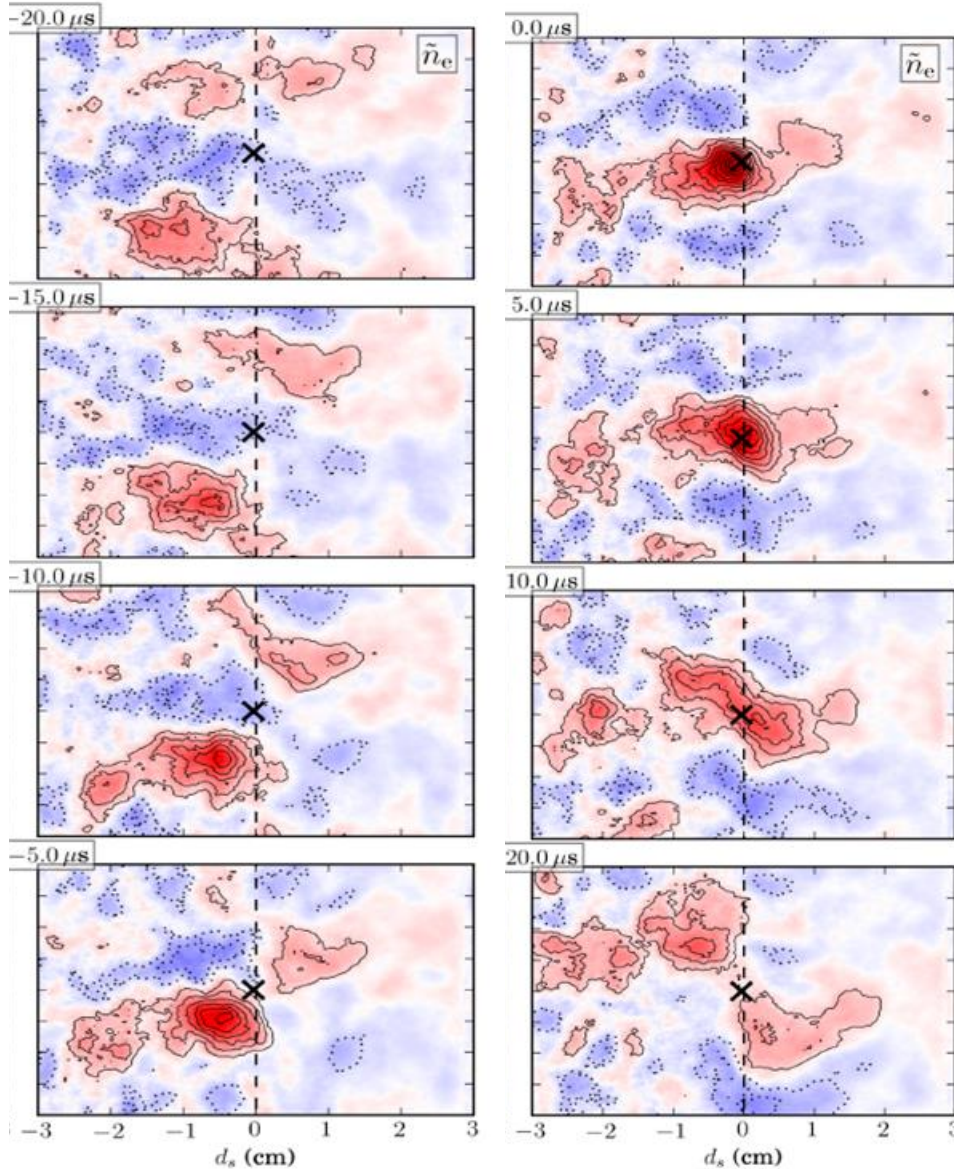
D.A D'Ippolito et al. Phys. Plasmas 2011

-Interchange instability generates radial extended structures (streamer)

-Shear flow breaks streamer up into blobs

- delta-f gyrofluid code
- electron and ion density (-> potential)
- perpendicular and parallel ion and electron temperature (finite Larmor radius effects)
- parallel ion dynamics (ion sound waves)
- electromagnetic (magnetic fluctuations)
- solves equilibrium (global)
- 3D
- transition from closed to open field lines (sheath connected boundary conditions) at separatrix





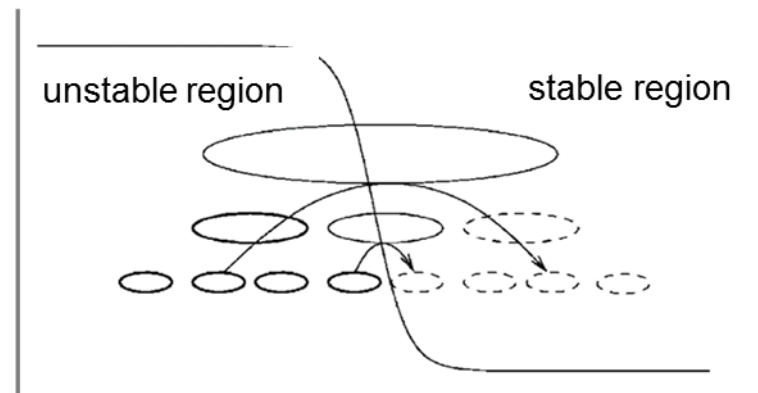
## Blob generation

- blobs merge
- exchange particles and energy
- break up

B. Nold et al. Phys. Plasmas 2014

(turbulence spreading)

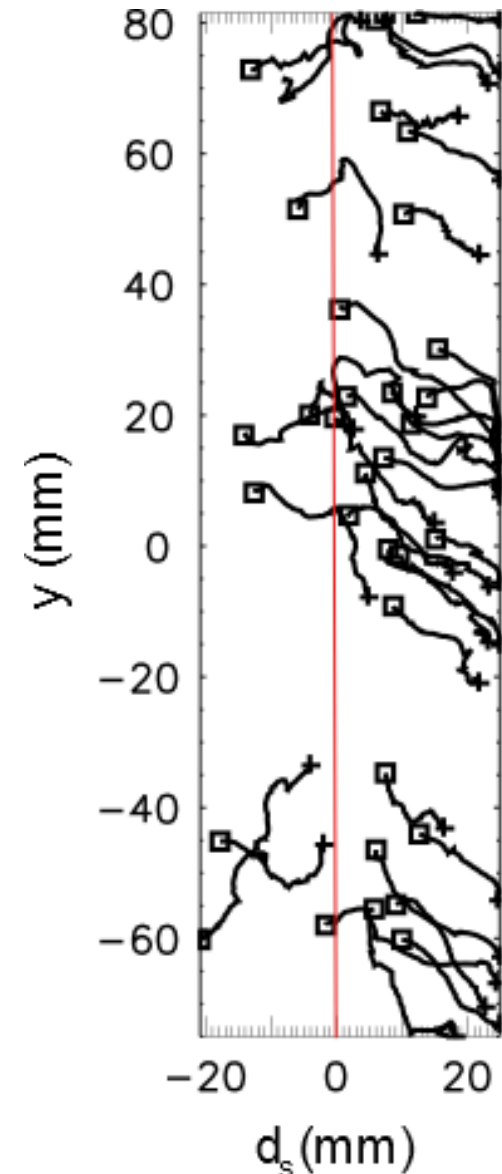
O.Gürçan et al. Phys. Plasmas 2005



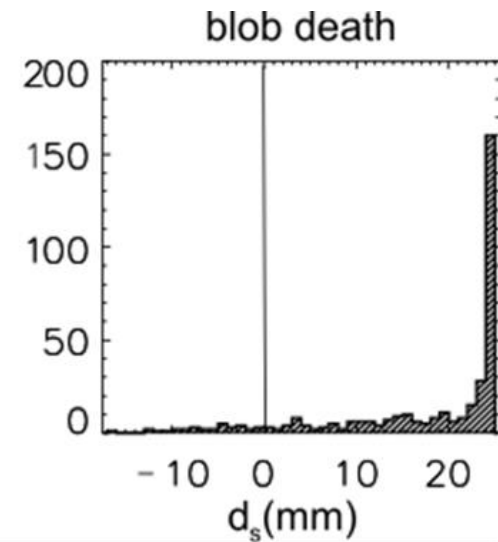
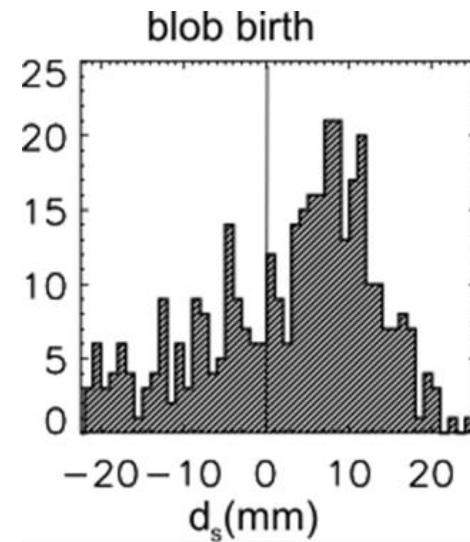
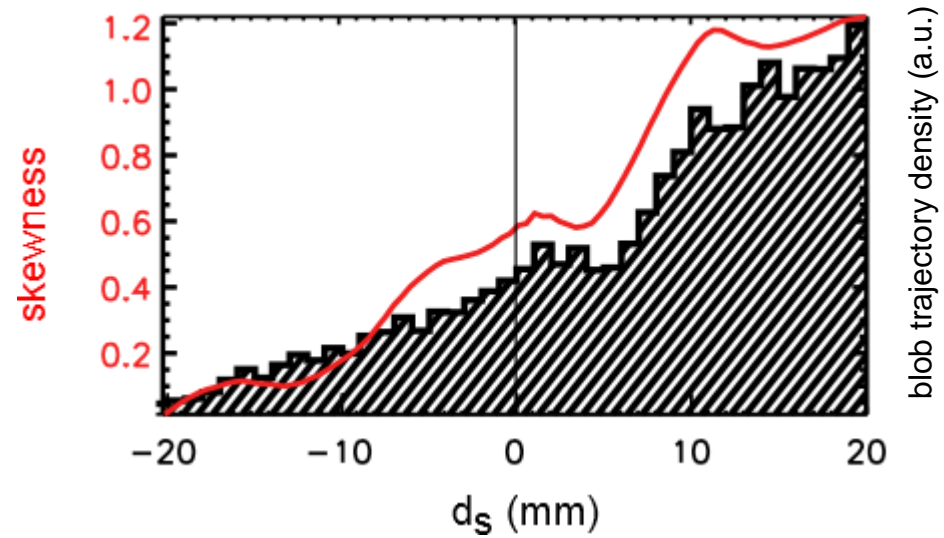
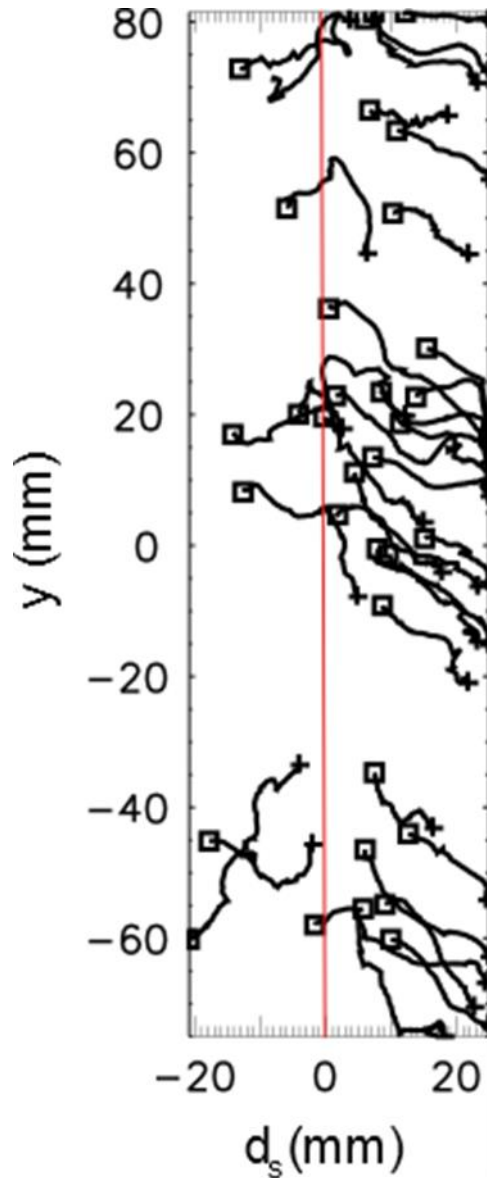
Blob definition:

Identify possible blobs at every time step:

- (i) positive density perturbations exceeding the standard deviation by a factor of 2.5
- (ii) has to fulfill (i) over a connected spatial extent  
(11 points in the simulation grid  $\sim 5\text{mm}$ )
- (iii) track possible blobs for a least 50 time steps ( $\sim 25\ \mu\text{s}$ )



# Statistics of blob trajectories



mean free energy

$$\frac{1}{2} \frac{\partial \langle n \rangle^2}{\partial t} = - \frac{\partial}{\partial r} (\langle n \rangle \langle \tilde{v}_r \tilde{n} \rangle) + \left\langle \frac{\partial n}{\partial r} \right\rangle \langle \tilde{v}_r \tilde{n} \rangle.$$

turbulent energy

$$\frac{1}{2} \frac{\partial \langle \tilde{n}^2 \rangle}{\partial t} = - \left\langle \frac{\partial n}{\partial r} \right\rangle \langle \tilde{v}_r \tilde{n} \rangle - \frac{1}{2} \frac{\partial}{\partial r} \langle \tilde{v}_r \tilde{n}^2 \rangle$$



mean free energy

$$\frac{1}{2} \frac{\partial \langle n \rangle^2}{\partial t} = - \frac{\partial}{\partial r} (\langle n \rangle \langle \tilde{v}_r \tilde{n} \rangle) + \left\langle \frac{\partial n}{\partial r} \right\rangle \langle \tilde{v}_r \tilde{n} \rangle.$$

turbulent energy

$$\frac{1}{2} \frac{\partial \langle \tilde{n}^2 \rangle}{\partial t} = - \left\langle \frac{\partial n}{\partial r} \right\rangle \langle \tilde{v}_r \tilde{n} \rangle - \frac{1}{2} \frac{\partial}{\partial r} \langle \tilde{v}_r \tilde{n}^2 \rangle$$

energy exchange



mean free energy

$$\frac{1}{2} \frac{\partial \langle n \rangle^2}{\partial t} = - \frac{\partial}{\partial r} (\langle n \rangle \langle \tilde{v}_r \tilde{n} \rangle) + \left\langle \frac{\partial n}{\partial r} \right\rangle \langle \tilde{v}_r \tilde{n} \rangle.$$

turbulent energy

$$\frac{1}{2} \frac{\partial \langle \tilde{n}^2 \rangle}{\partial t} = - \left\langle \frac{\partial n}{\partial r} \right\rangle \langle \tilde{v}_r \tilde{n} \rangle - \frac{1}{2} \frac{\partial}{\partial r} \langle \tilde{v}_r \tilde{n}^2 \rangle$$

energy exchange

local drive

turbulent spreading

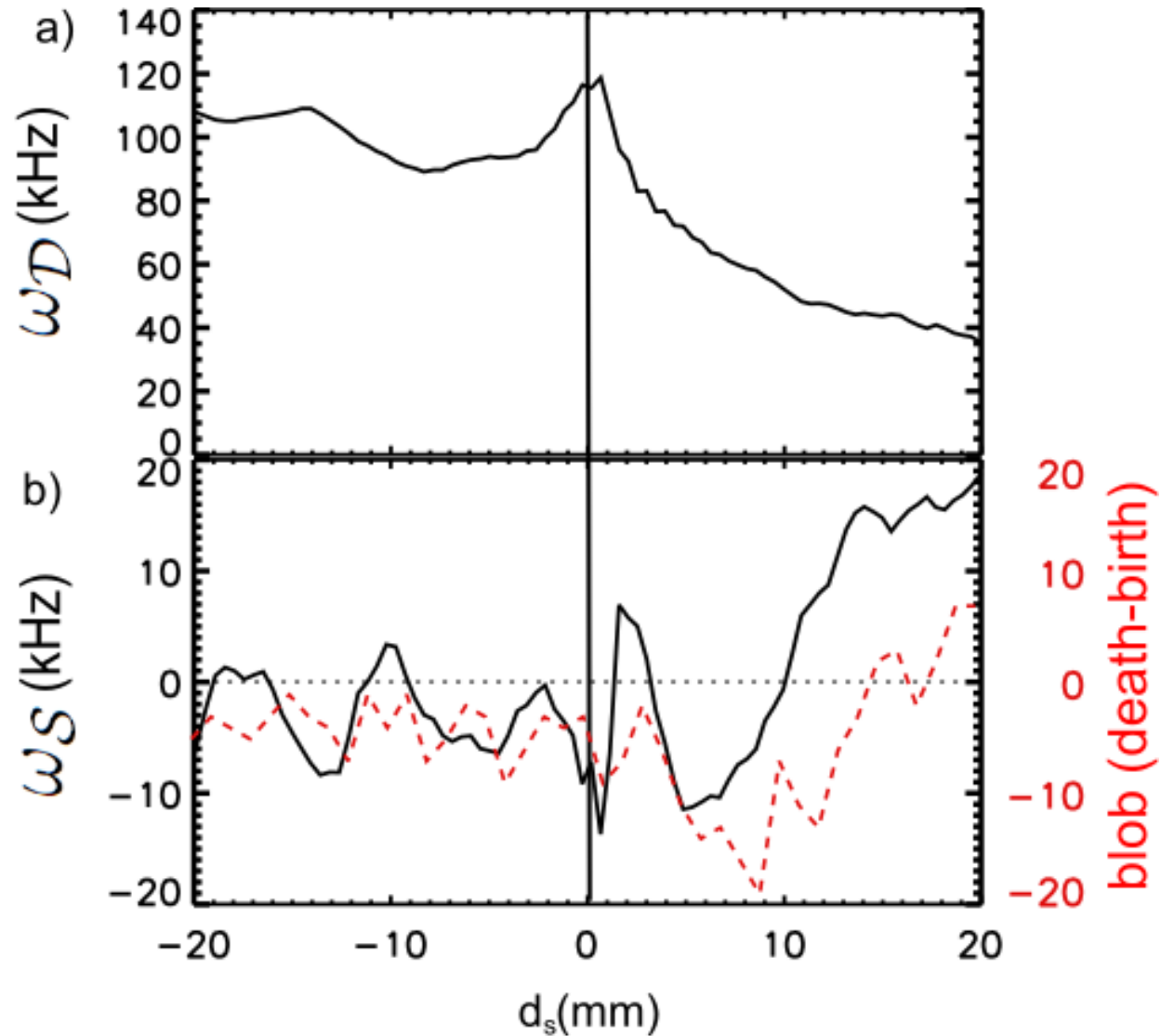
$$\omega_{\mathcal{D}} = \frac{- \left\langle \frac{\partial n}{\partial r} \right\rangle \langle \tilde{v}_r \tilde{n} \rangle}{\frac{1}{2} \langle \tilde{n}^2 \rangle}$$

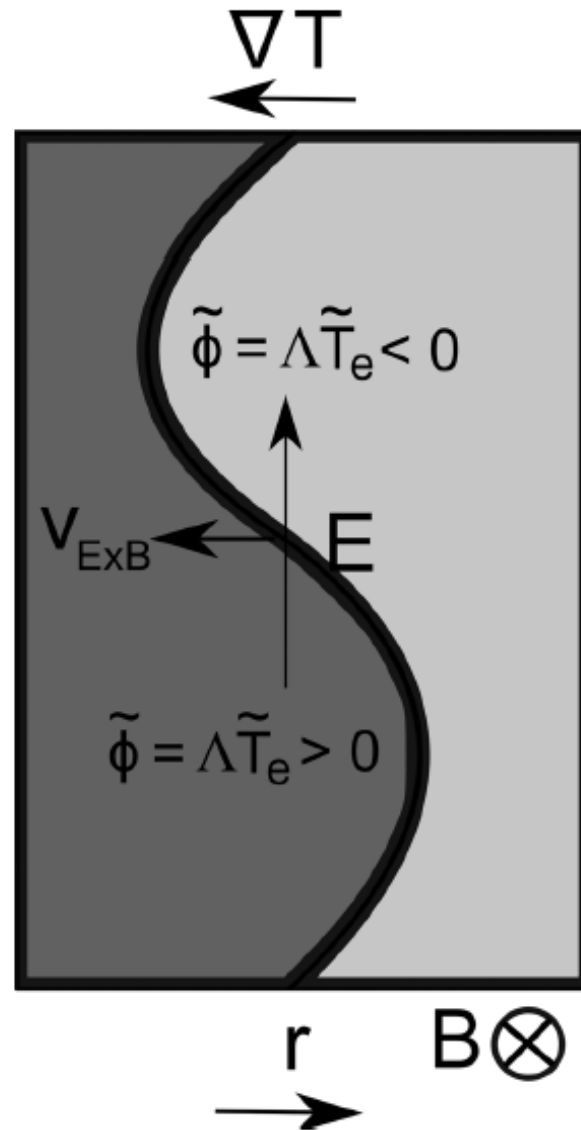
$$\omega_{\mathcal{S}} = \frac{- \frac{1}{2} \frac{\partial}{\partial r} \langle \tilde{v}_r \tilde{n}^2 \rangle}{\frac{1}{2} \langle \tilde{n}^2 \rangle}$$

# Local drive vs. spreading

$$\omega_{\mathcal{D}} = \frac{-\langle \frac{\partial n}{\partial r} \rangle \langle \tilde{v}_r \tilde{n} \rangle}{\frac{1}{2} \langle \tilde{n}^2 \rangle}$$

$$\omega_{\mathcal{S}} = \frac{-\frac{1}{2} \frac{\partial}{\partial r} \langle \tilde{v}_r \tilde{n}^2 \rangle}{\frac{1}{2} \langle \tilde{n}^2 \rangle}$$



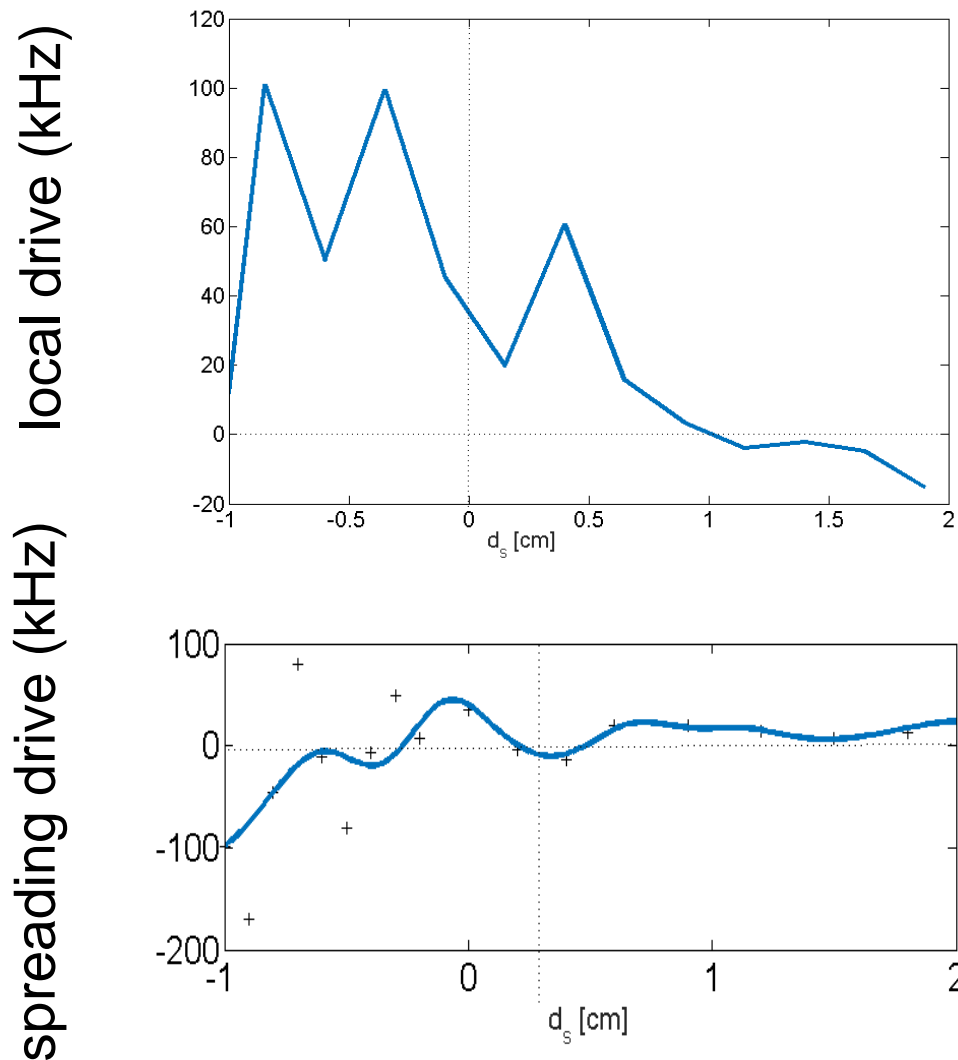
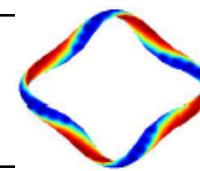


see talk by  
Valentina Nikolaeva  
today

P. Manz et al. Phys. Plasmas 2015

# Results from HSX

ETDB ([www.ipp.mpg.de/ISS/](http://www.ipp.mpg.de/ISS/))



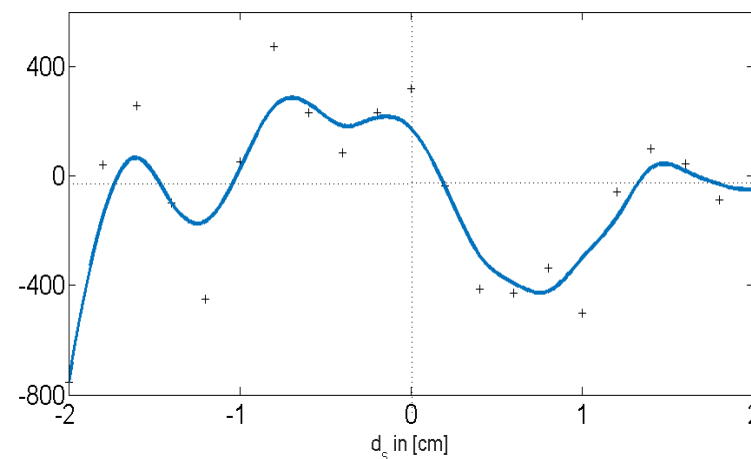
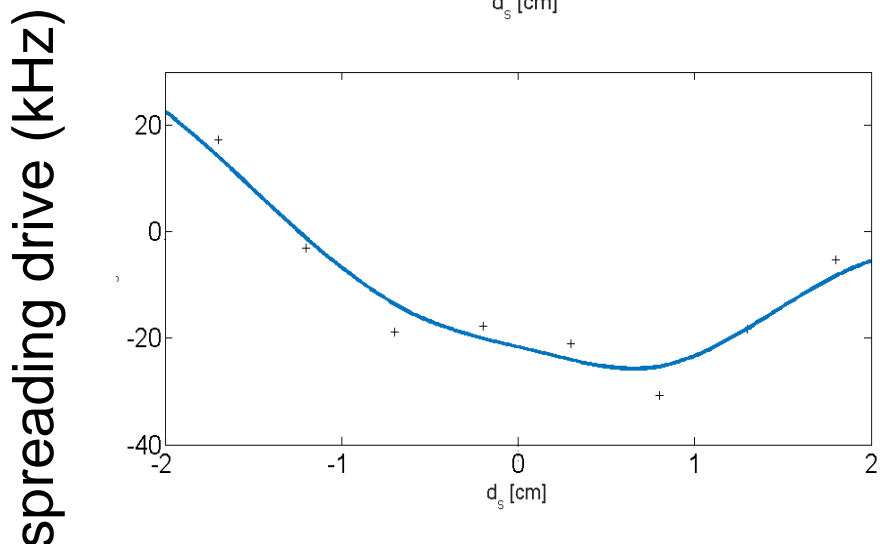
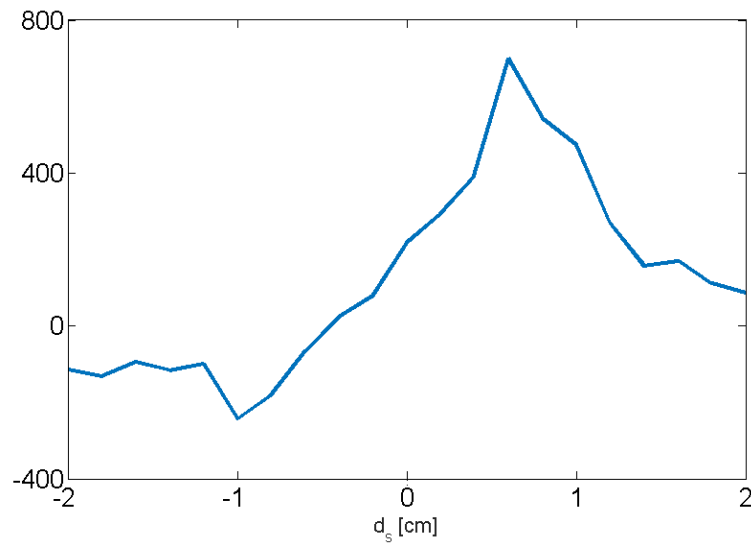
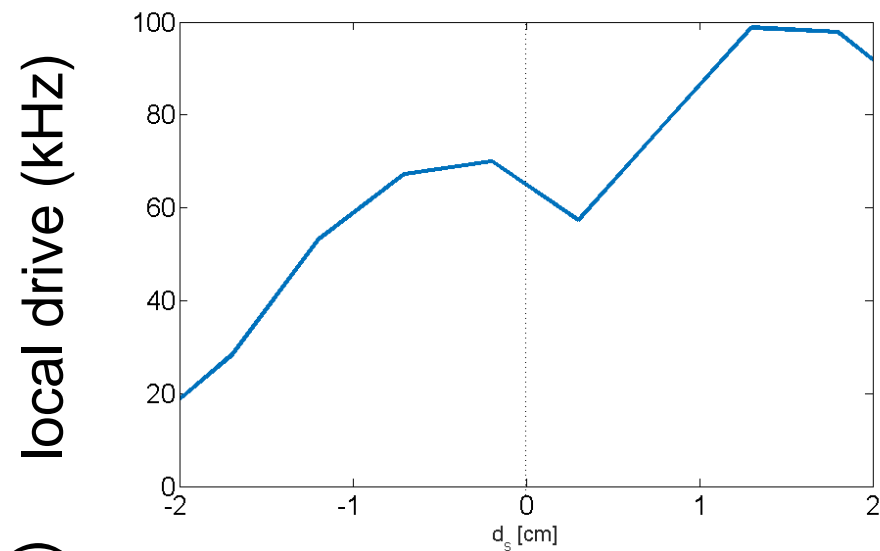
# Results from TJ-K and WEGA

ETDB ([www.ipp.mpg.de/ISS/](http://www.ipp.mpg.de/ISS/))



## TJ-K

## WEGA



- Blobs are not generated at one particular position
- At ASDEX Upgrade L-mode sheath connected conditions most blobs are generated outside the separatrix (CWI)
- Turbulence spreading a more suitable diagnostics for blob generation than the skewness alone
- Turbulence spreading should play the key role for turbulence in the far-SOL (once the background gradient is small)

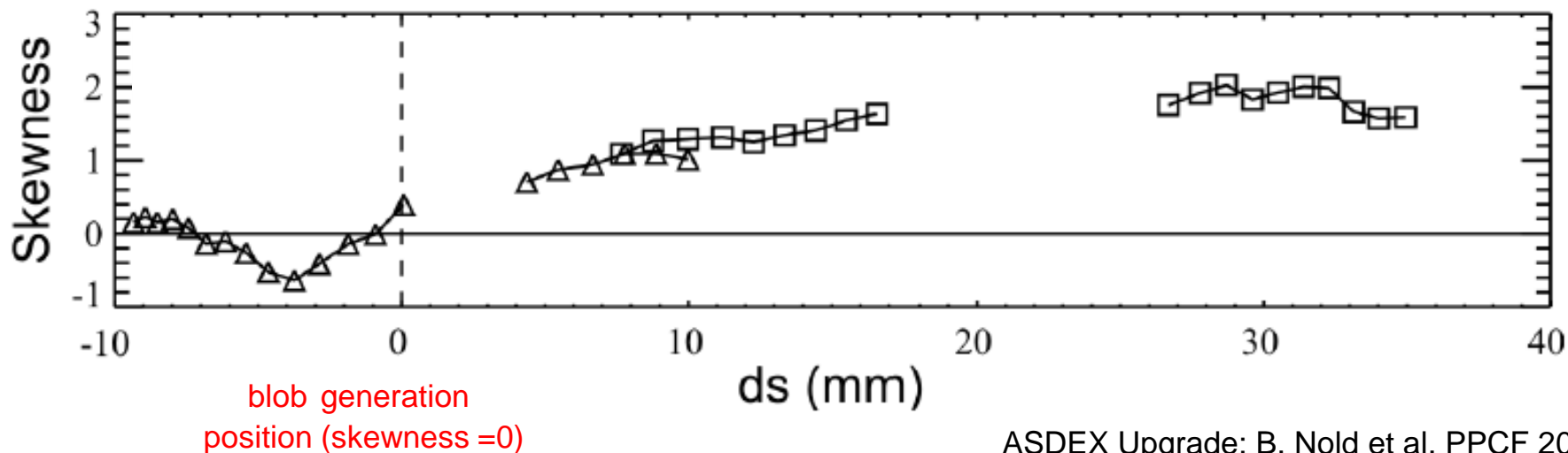
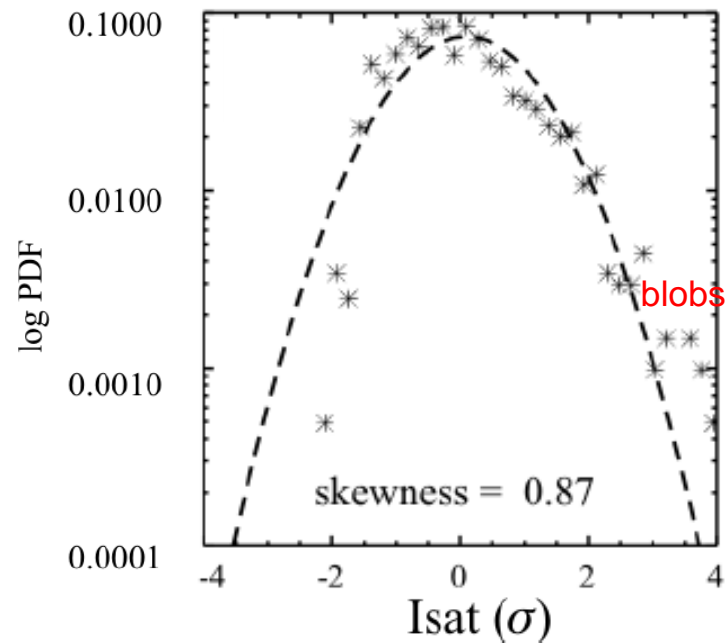
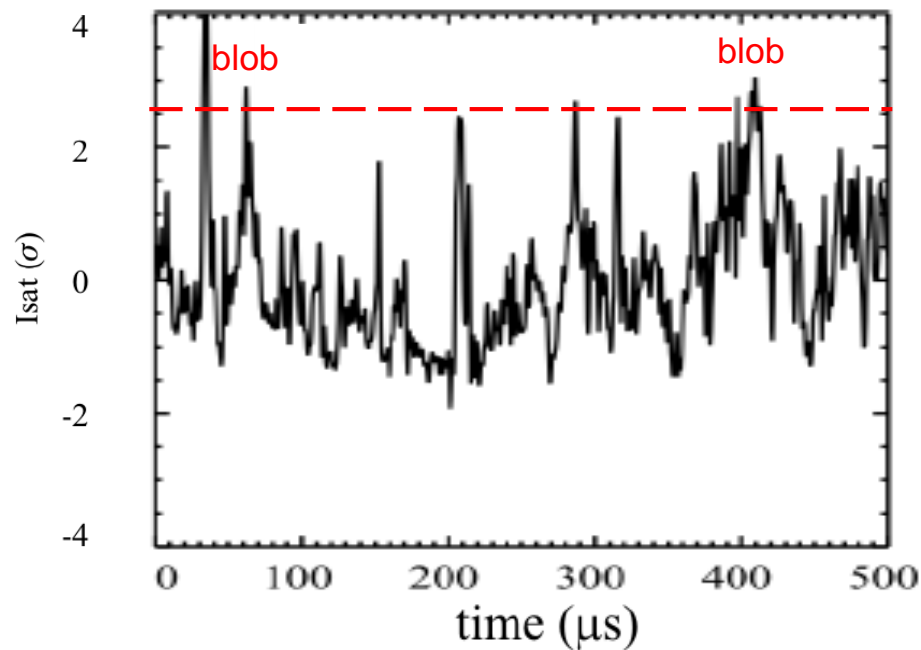
local drive

$$\omega_{\mathcal{D}} = \frac{-\langle \frac{\partial n}{\partial r} \rangle \langle \tilde{v}_r \tilde{n} \rangle}{\frac{1}{2} \langle \tilde{n}^2 \rangle}$$

turbulent spreading

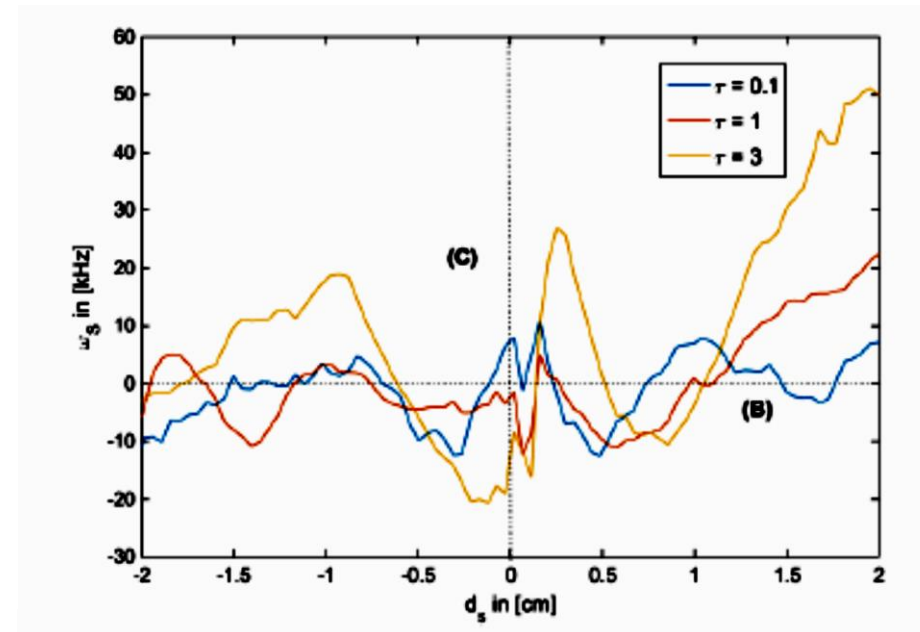
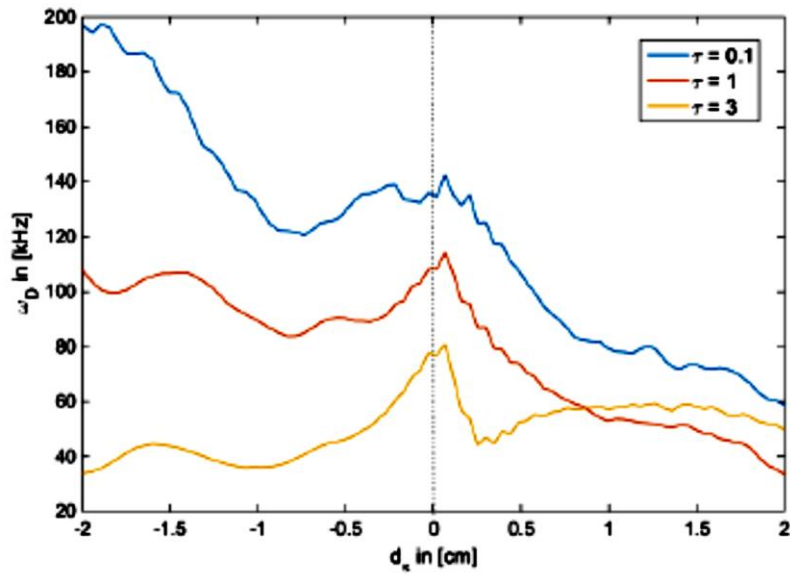
$$\omega_{\mathcal{S}} = \frac{-\frac{1}{2} \frac{\partial}{\partial r} \langle \tilde{v}_r \tilde{n}^2 \rangle}{\frac{1}{2} \langle \tilde{n}^2 \rangle}$$

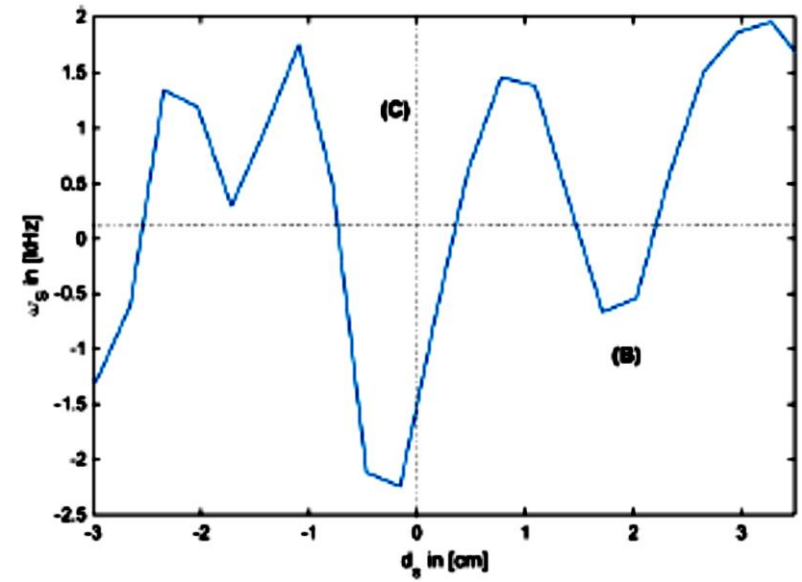
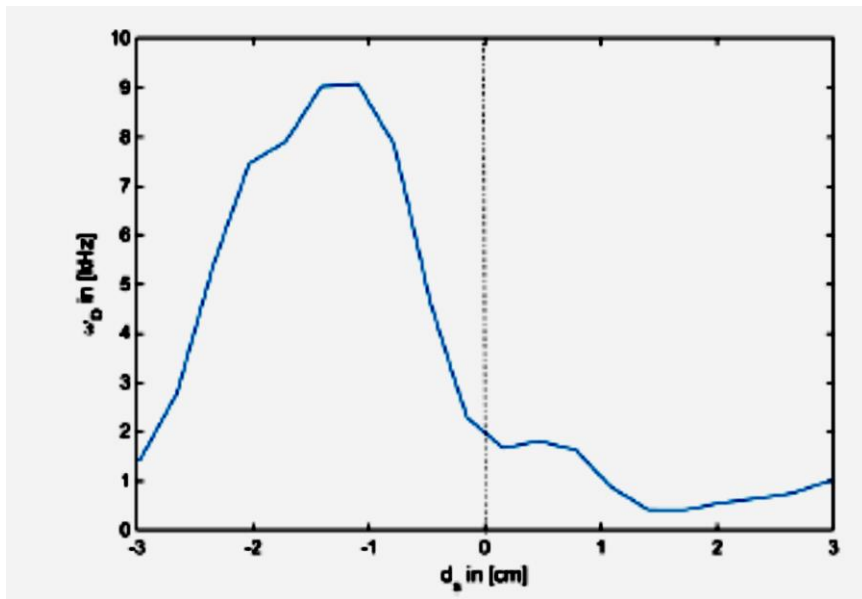
# Skewness as the basic parameter



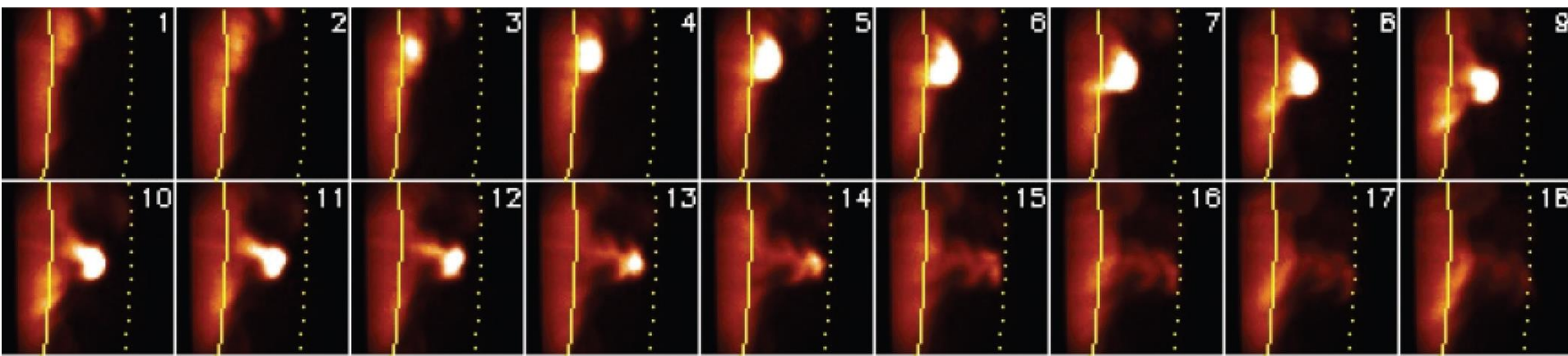


# Additional material





# Blobs propagate radially outward



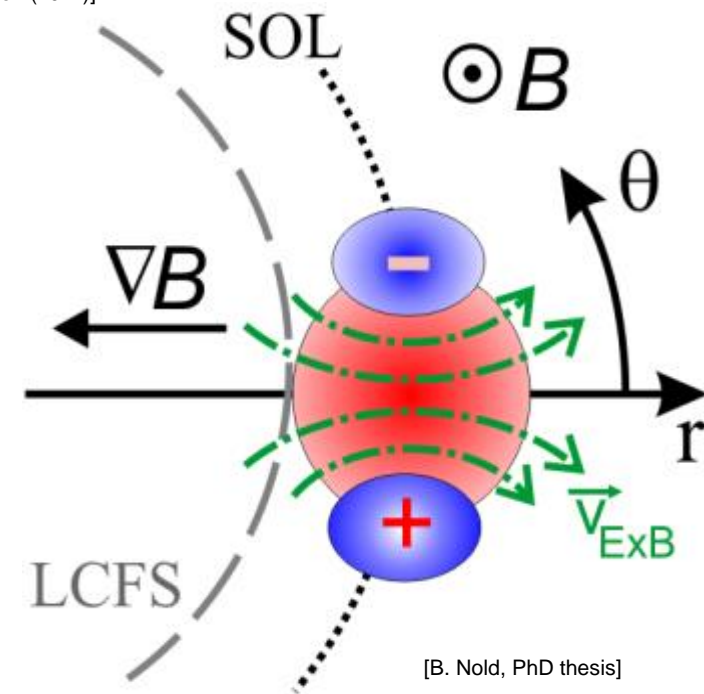
[GPI by R. J. Maqueda, D. P. Stotler, S. J. Zweben and The NSTX team, J. Nucl. Mater. (2011)]

charge separation by curvature/gradB drift

$$\mathbf{v}_D^{\nabla B} = -\frac{W_{\perp}}{q} \frac{\nabla_{\perp} B \times \mathbf{B}}{B^3}$$

ExB drift accelerates blobs radially

$$\mathbf{v}_D^{E \times B} = \frac{\mathbf{E} \times \mathbf{B}}{B^2}$$



[B. Nold, PhD thesis]

**Blobs are intermittently expelled density filaments in the scrape-off layer (SOL)**

- Elongated along the magnetic field line
- Low magnetic field component (in contrast to ELMs)

