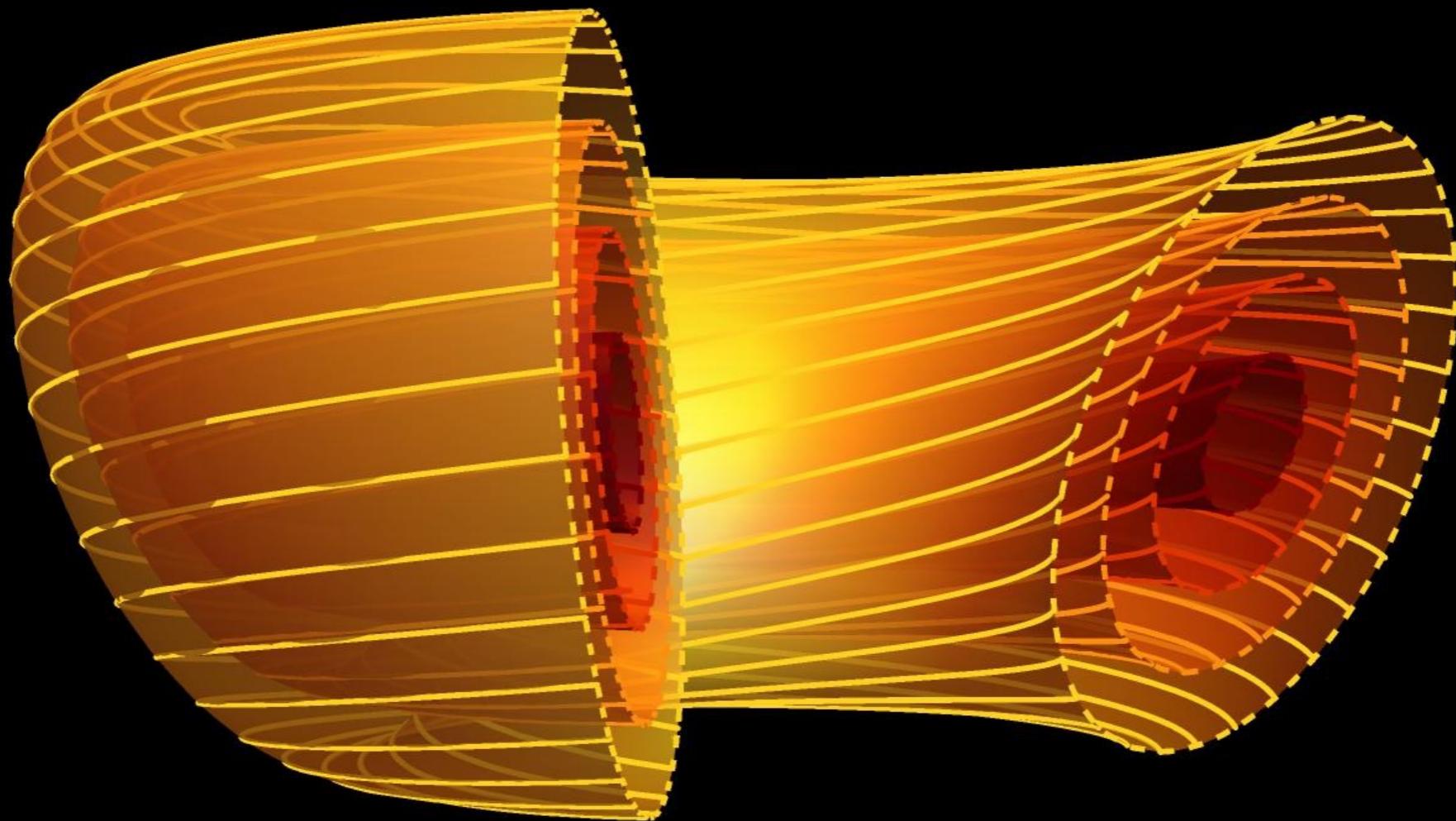


# Intrinsic momentum transport in tokamaks with tilted elliptical flux surfaces



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European Fusion Theory Conference  
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# The problem

Troyon et al. *PPCF* (1984).  
Liu et al. *Nucl. Fusion* (2004).  
Garofalo et al. *PRL* (2002).

- Make tokamaks work better
- Do this by increasing the limit on  $\beta$ :  $\beta_N \equiv \frac{aB_0}{I_p} \frac{2\mu_0 \langle p \rangle}{B_0^2} \lesssim \beta_N^{\text{Troyon}}$
- Nearly all reactor-scale devices excite instabilities by violating the  $\beta$  limit, e.g.

$$\beta_N \approx 2.57 > \beta_N^{\text{Troyon}} \approx 2.45 \quad (\text{ITER})$$

- Stabilizing ITER at  $\beta_N = 3$  through toroidal rotation requires

$$M_A \equiv \frac{R\Omega_\zeta}{v_{\text{Alfvén}}} \approx 1 - 5\% \quad \text{*} \quad (\text{on-axis})$$

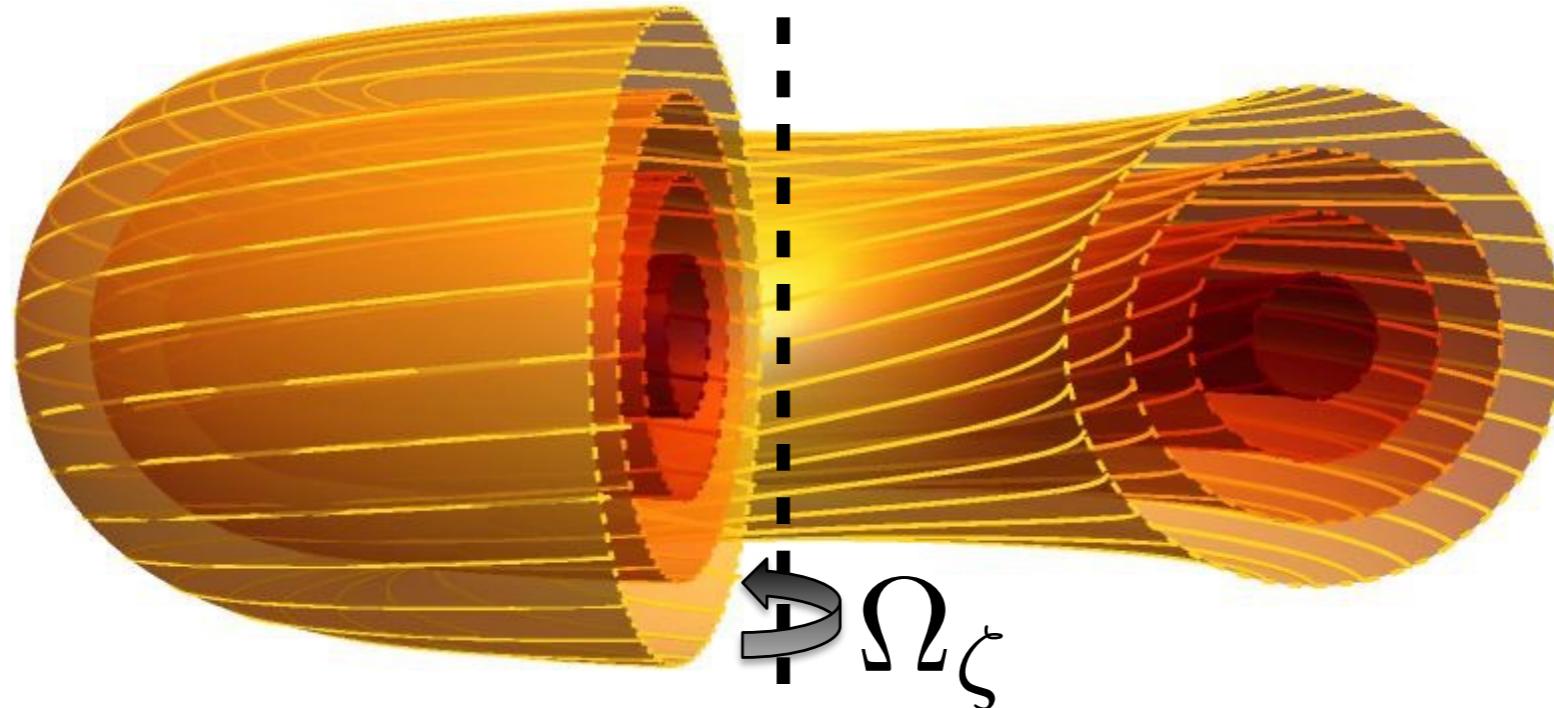
\* multiply by  $\sim 10$  to get

$$M_S \equiv \frac{R\Omega_\zeta}{v_{\text{sound}}}$$

- DIII-D used  $\sim 4\%$  rotation to enable discharges with  $\beta_N \approx 1.5\beta_N^{\text{Troyon}}$

# Toroidal rotation

Liu et al. *Nucl. Fusion* (2004).



- Each flux surface is free to rotate toroidally because of axisymmetry
- A plasma will start to spin if you push it (neutral beams, RF waves) or if it pushes off itself (“intrinsic rotation” from turbulence)
- Rotation from neutral beams in ITER has a Alfvén Mach number of  $\sim 0.3\%$

# Intrinsic rotation

---

- Generated by the turbulent transport of momentum between flux surfaces
- Scales well to large devices and requires no external power
- Solve  $\left\langle \Pi \left( \Omega_\zeta, \frac{d\Omega_\zeta}{dr_\psi} \right) \right\rangle_t = 0$  for steady state  $\Omega_\zeta(r_\psi)$ :  
$$\left\langle \Pi \left( \Omega_\zeta = 0, \frac{d\Omega_\zeta}{dr_\psi} = 0 \right) \right\rangle_t - P_\Pi \Omega_\zeta - D_\Pi \frac{d\Omega_\zeta}{dr_\psi} \approx 0$$
- All we need is for the turbulent transport of momentum in the absence of both rotation and rotation shear **to be non-zero**, i.e

$$\langle \Pi(0, 0) \rangle_t \neq 0$$

“The turbulent transport of momentum in  
the absence of both rotation and rotation  
shear **must be zero**...”

Peeters, et al. *Phys. Plasmas* (2005).

Parra, et al. *Phys. Plasmas* (2011).

Sugama, et al. *PPCF* (2011).

“The turbulent transport of momentum in  
the absence of both rotation and rotation  
shear must be zero,  
**for up-down symmetric tokamaks.”**

Peeters, et al. *Phys. Plasmas* (2005).

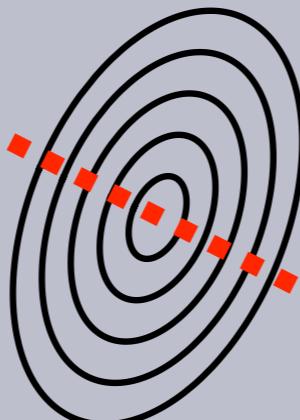
Parra, et al. *Phys. Plasmas* (2011).

Sugama, et al. *PPCF* (2011).

# Outline

## Tokamaks

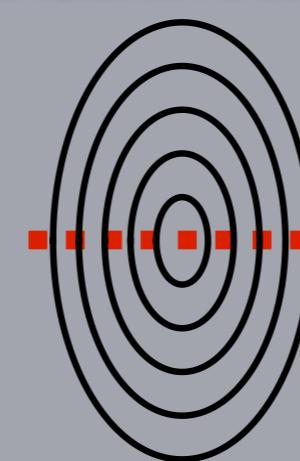
Mirror  
symmetric



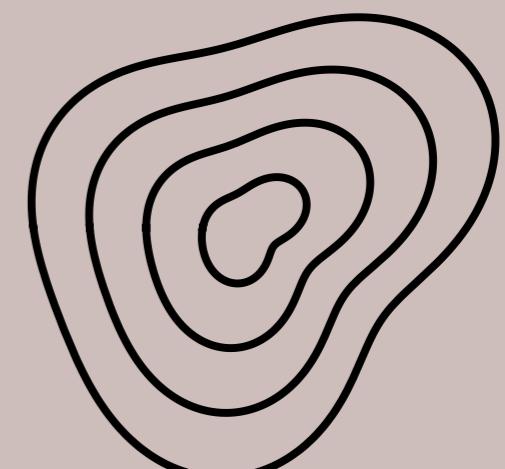
Up-down  
symmetric

$$M_A = 0$$

$$\langle \Pi(0, 0) \rangle_t = 0$$



Non-mirror  
symmetric



# Gyrokinetics

---

- Governs turbulence in tokamaks:

$$\begin{aligned} \frac{\partial h_s}{\partial t} + v_{||} \hat{b} \cdot \vec{\nabla} \theta \left. \frac{\partial h_s}{\partial \theta} \right|_{v_{||}} &+ i (k_\psi v_{ds\psi} + k_\alpha v_{ds\alpha}) h_s + a_{||s} \frac{\partial h_s}{\partial v_{||}} - \sum_{s'} \langle C_{ss'}^{(l)} \rangle_\varphi + \{ J_0 (k_\perp \rho_s) \phi, h_s \} \\ &= \frac{Z_s e F_{Ms}}{T_s} \frac{\partial}{\partial t} (J_0 (k_\perp \rho_s) \phi) - v_{\phi s\psi} F_{Ms} \left[ \frac{1}{n_s} \frac{dn_s}{d\psi} + \left( \frac{m_s v^2}{2T_s} - \frac{3}{2} \right) \frac{1}{T_s} \frac{dT_s}{d\psi} \right] \end{aligned}$$

where  $k_\perp = \sqrt{k_\psi^2 \left| \vec{\nabla} \psi \right|^2 + 2k_\psi k_\alpha \vec{\nabla} \psi \cdot \vec{\nabla} \alpha + k_\alpha^2 \left| \vec{\nabla} \alpha \right|^2}$

- Allows us to calculate the momentum flux

$$\Pi = 2\pi i I \sum_{k_\psi, k_\alpha} k_\alpha \left\langle \phi (k_\psi, k_\alpha) \int dv_{||} d\mu v_{||} J_0 (k_\perp \rho_s) h_s (-k_\psi, -k_\alpha) \right\rangle_\psi$$

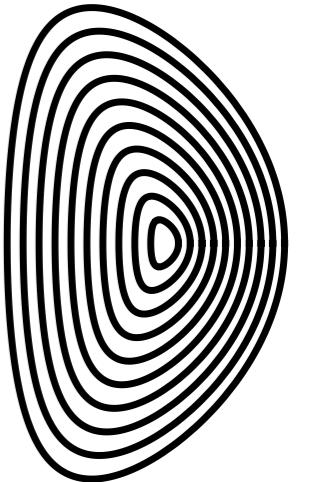
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$$\Pi = 2\pi i I \sum_{k_\psi, k_\alpha} k_\alpha \left\langle \phi(k_\psi, k_\alpha) \int dv_{||} d\mu v_{||} J_0 (k_\perp \rho_s) h_s (-k_\psi, -k_\alpha) \right\rangle_\psi$$

- Calculate the eight geometric coefficients from MHD equilibrium

# Generalization of Miller local equilibrium

Miller et al. *Phys. Plasmas* (1998).

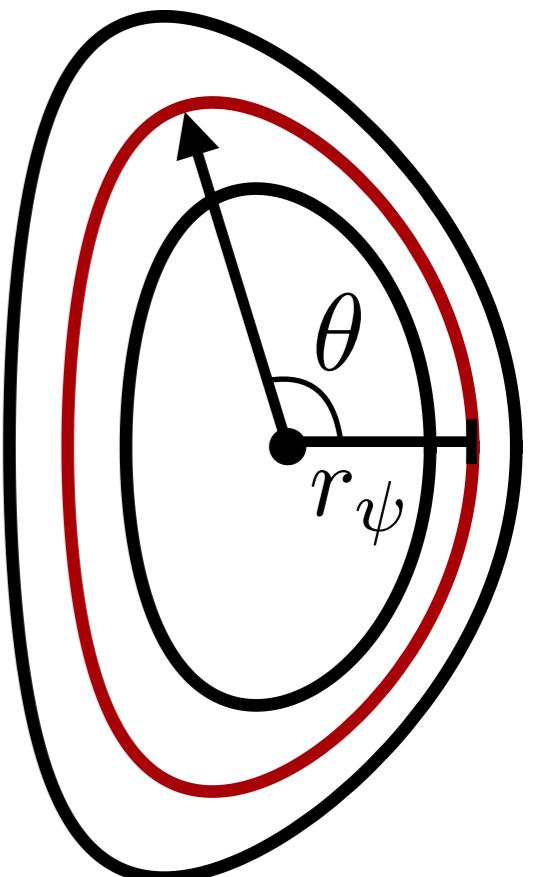
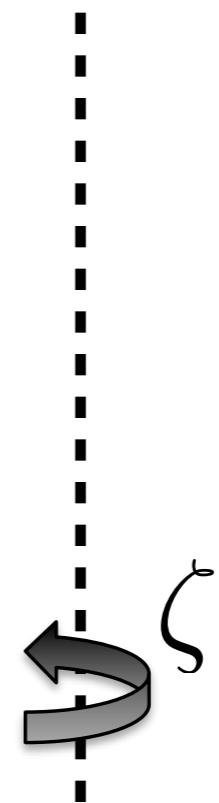
- Works well with GS2, a local  $\delta f$  gyrokinetic code

- Specify the flux surface of interest as a Fourier decomposition:

$$r_0(r_\psi, \theta) = r_\psi \left( 1 - \sum_m C_m \cos(m(\theta + \theta_{tm})) \right)$$

- Specify how it changes with minor radius:

$$\frac{\partial r_0}{\partial r_\psi} \Big|_\theta = 1 - \sum_m C'_m \cos(m(\theta + \theta'_{tm}))$$



# Rotation in up-down symmetric tokamaks

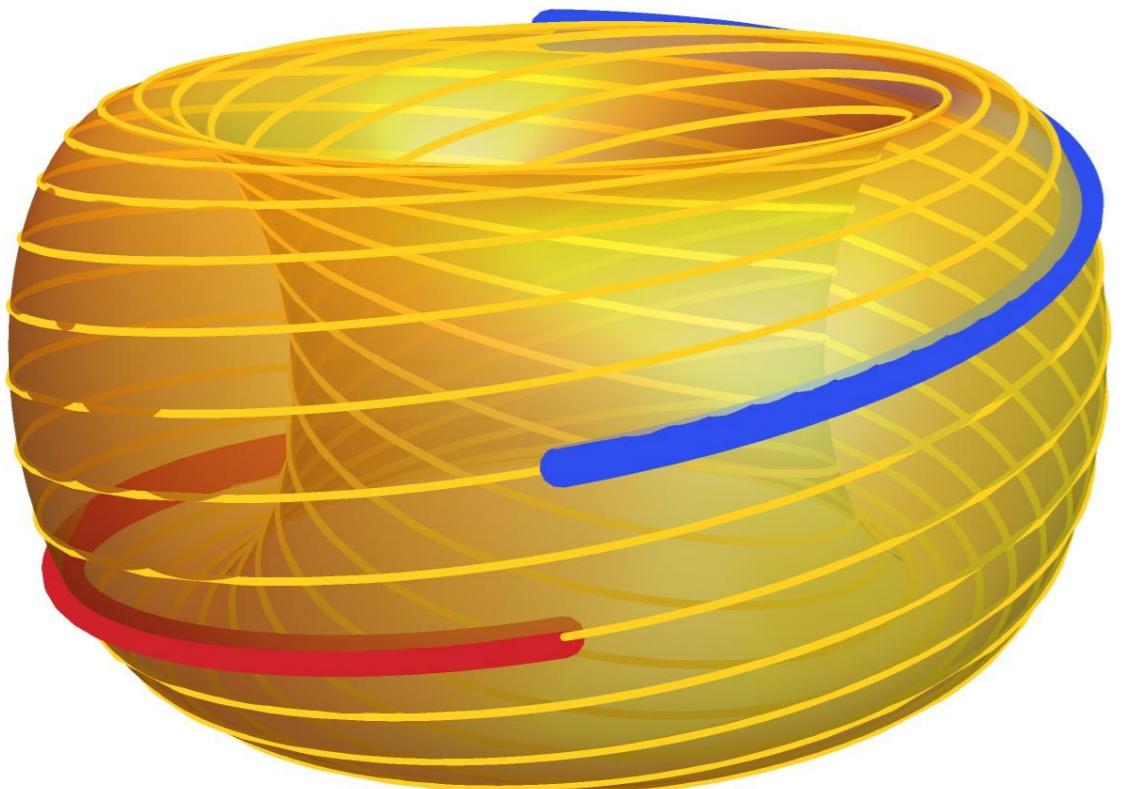
Peeters et al. *PoP* (2005). & Parra et al. *PoP* (2011).

Sugama et al. *PPCF* (2011).

- Negating  $k_\psi$ ,  $\theta$ , and  $v_{||}$  leads to a second solution of the gyrokinetic eq.

$$\begin{aligned} \left\{ B, \hat{b} \cdot \vec{\nabla} \theta, v_{ds\psi}, v_{ds\alpha}, a_{||s}, |\vec{\nabla} \psi|^2, \vec{\nabla} \psi \cdot \vec{\nabla} \alpha, |\vec{\nabla} \alpha|^2 \right\} \\ \rightarrow \left\{ B, \hat{b} \cdot \vec{\nabla} \theta, -v_{ds\psi}, v_{ds\alpha}, -a_{||s}, |\vec{\nabla} \psi|^2, -\vec{\nabla} \psi \cdot \vec{\nabla} \alpha, |\vec{\nabla} \alpha|^2 \right\} \end{aligned}$$

$$h_s(k_\psi, k_\alpha, \theta, v_{||}, \mu, t) \rightarrow -h_s(-k_\psi, k_\alpha, -\theta, -v_{||}, \mu, t)$$



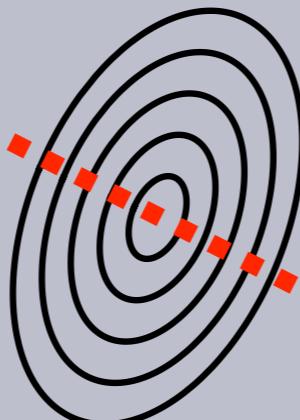
$$\langle \Pi(0,0) \rangle_t \rightarrow -\langle \Pi(0,0) \rangle_t$$

- Contributions to momentum flux from two symmetric particle trajectories cancel
- Constrains  $\langle \Pi(0,0) \rangle_t = 0$  to lowest order in  $\rho_* \equiv \rho_i/a \ll 1$

# Outline

## Tokamaks

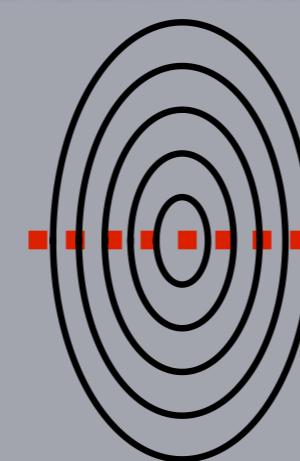
Mirror  
symmetric



Up-down  
symmetric

$$M_A = 0$$

$$\langle \Pi(0, 0) \rangle_t = 0$$



Non-mirror  
symmetric

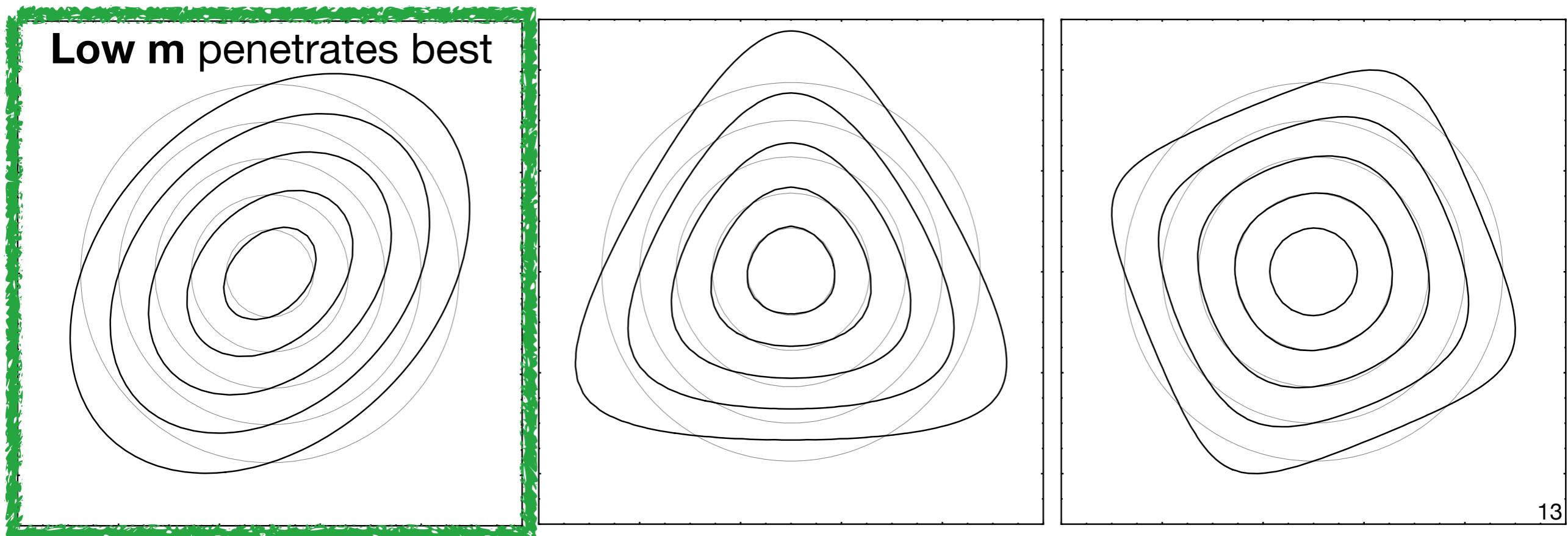


# MHD equilibrium of up-down asymmetry

Rodrigues et al. *Nucl. Fusion* (2014).

Ball et al. *PPCF* (2015).

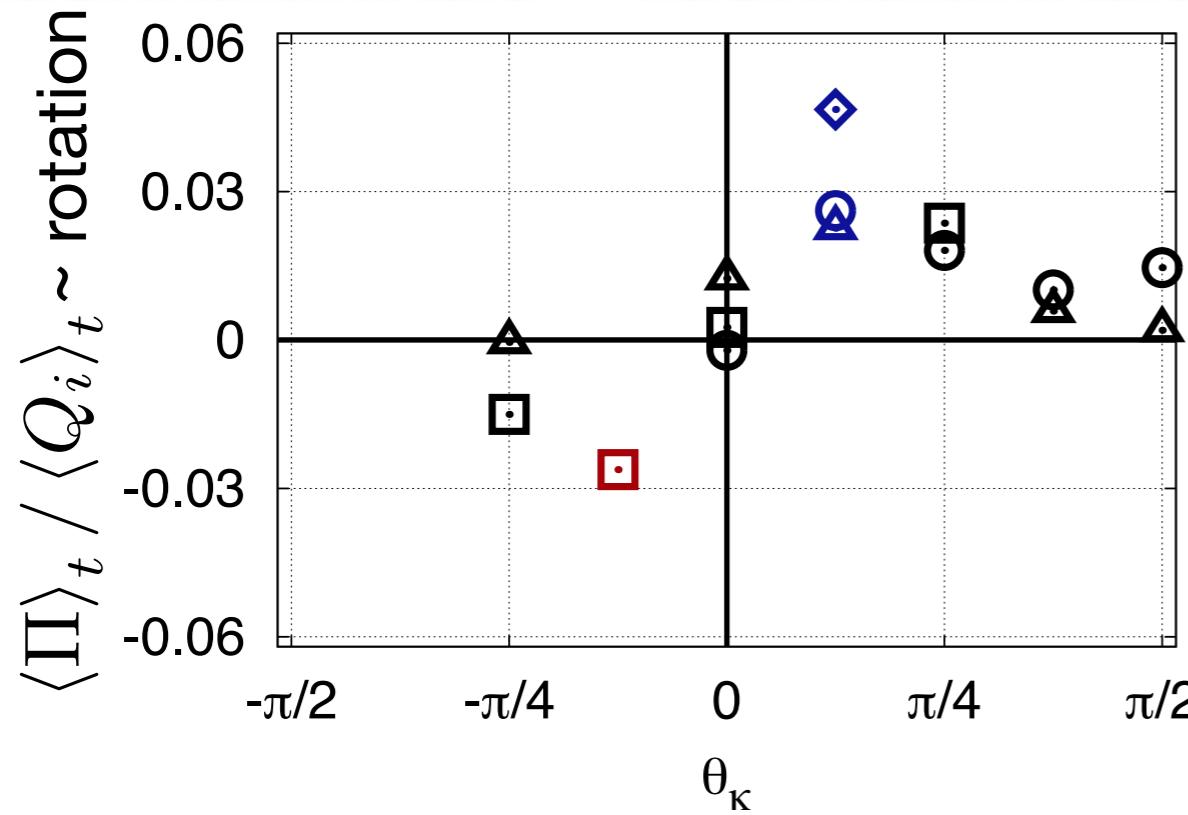
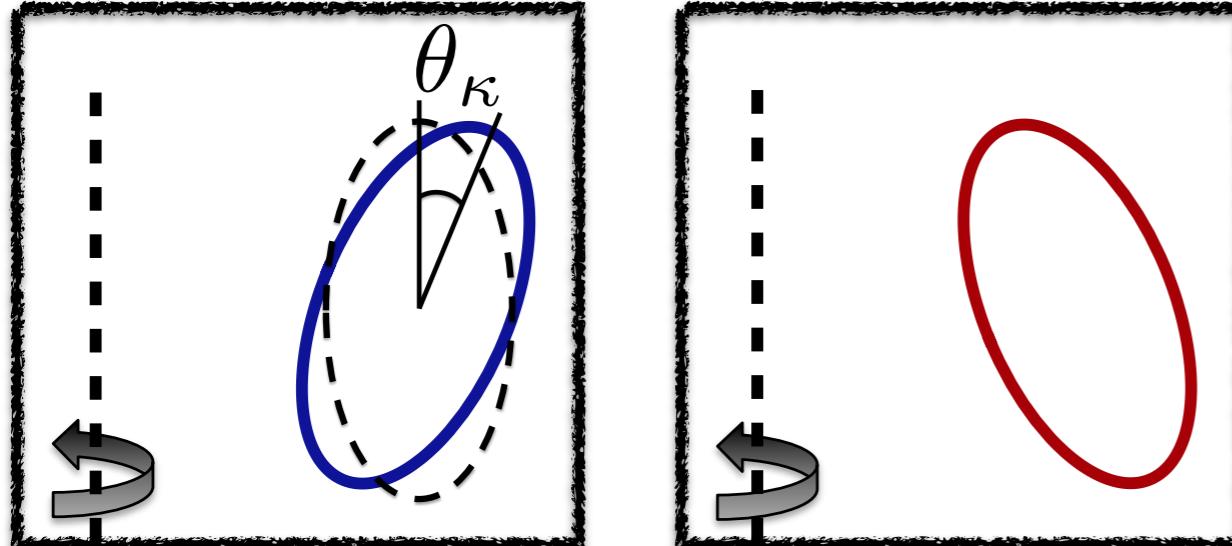
- Grad-Shafranov equation:  $R^2 \vec{\nabla} \cdot \left( \frac{\vec{\nabla} \psi}{R^2} \right) = -\mu_0 R^2 \frac{dp}{d\psi} - I \frac{dI}{d\psi}$
- Let the toroidal current be constant, i.e.  $-\mu_0 R^2 \frac{dp}{d\psi} - I \frac{dI}{d\psi} = c_0$
- To lowest order in aspect ratio, solutions are cylindrical harmonics:
  - m=2 mode
  - m=3 mode
  - m=4 mode



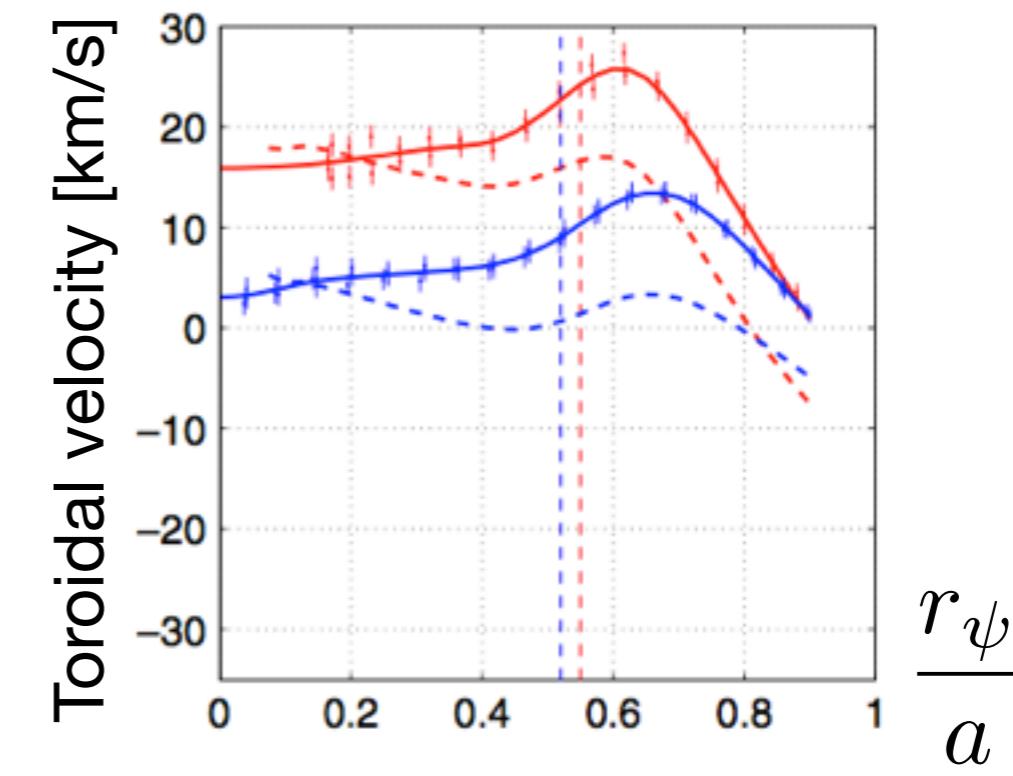
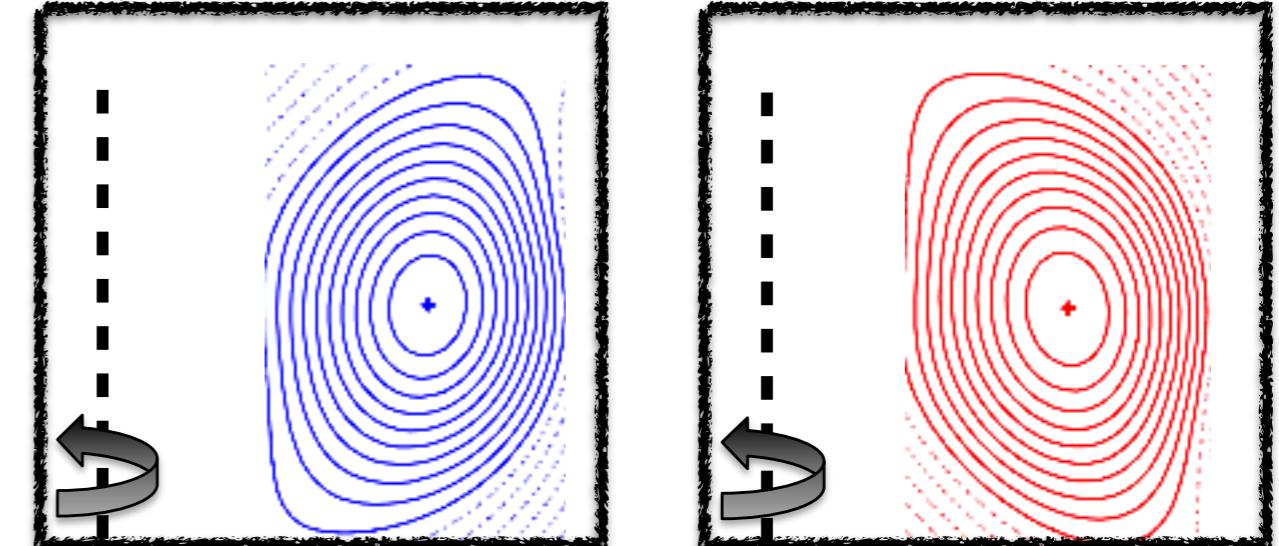
# Tilted elliptical geometry ( $m=2$ )

Ball et al. *PPCF* (2014).  
 Camenen et al. *PPCF* (2010).

## GS2 gyrokinetic simulations



## TCV experiment



# Estimating the Alfvén Mach number

Ball et al. *PPCF* (2014).  
Peeters et al. *PRL* (2007).

$$\langle \Pi(0,0) \rangle_t - P_\Pi \Omega_\zeta \xrightarrow{\text{ignore}} - D_\Pi \frac{d\Omega_\zeta}{dr_\psi} \approx 0$$

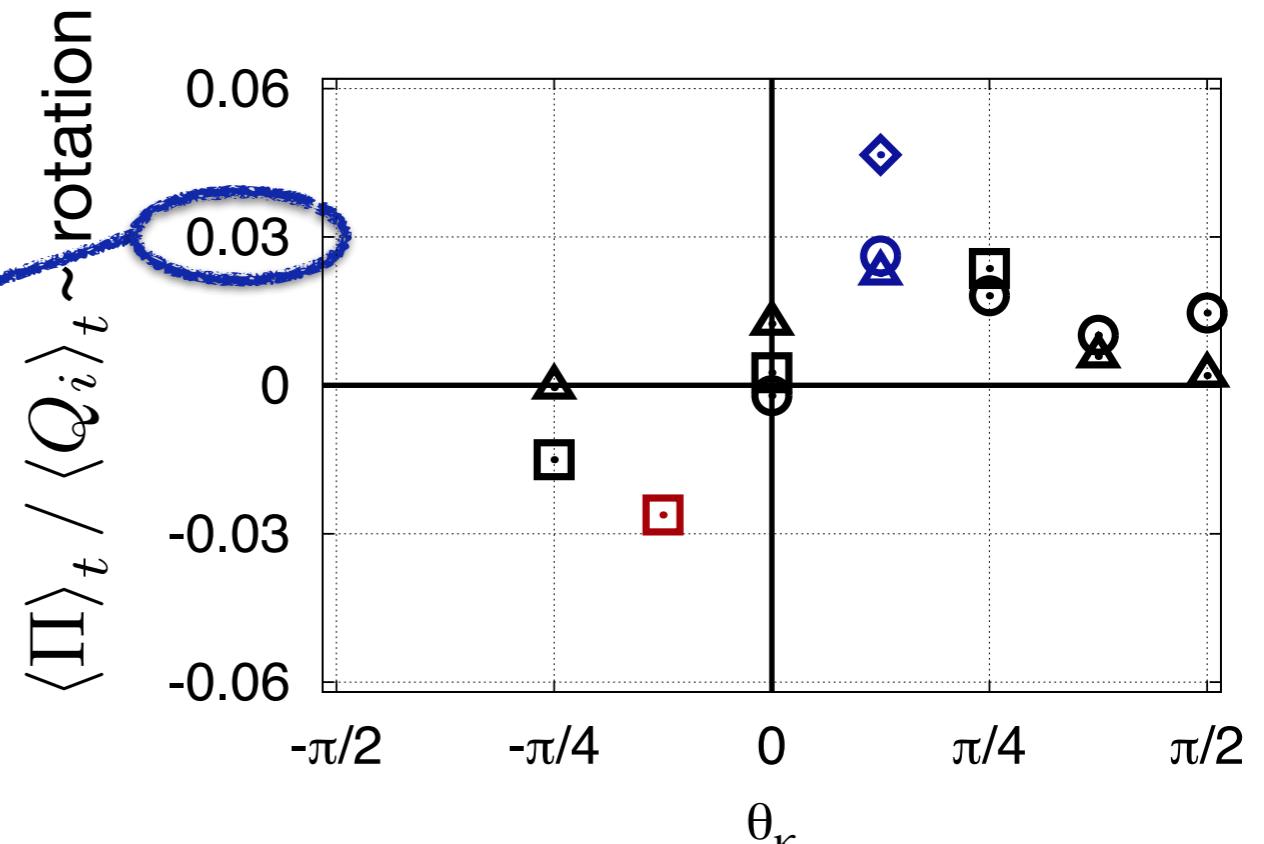
$$\langle Q_i \rangle_t \approx -D_Q \frac{dT_i}{dr_\psi}$$

$$Pr \equiv \frac{D_\Pi}{D_Q} \approx 1 \approx \text{constant}$$

$$M_A \approx \frac{\sqrt{2\beta_T}}{Pr} \frac{\langle \Pi(0,0) \rangle_t}{\langle Q_i \rangle_t}$$

$M_A \approx 1.5\%$

(in ITER)

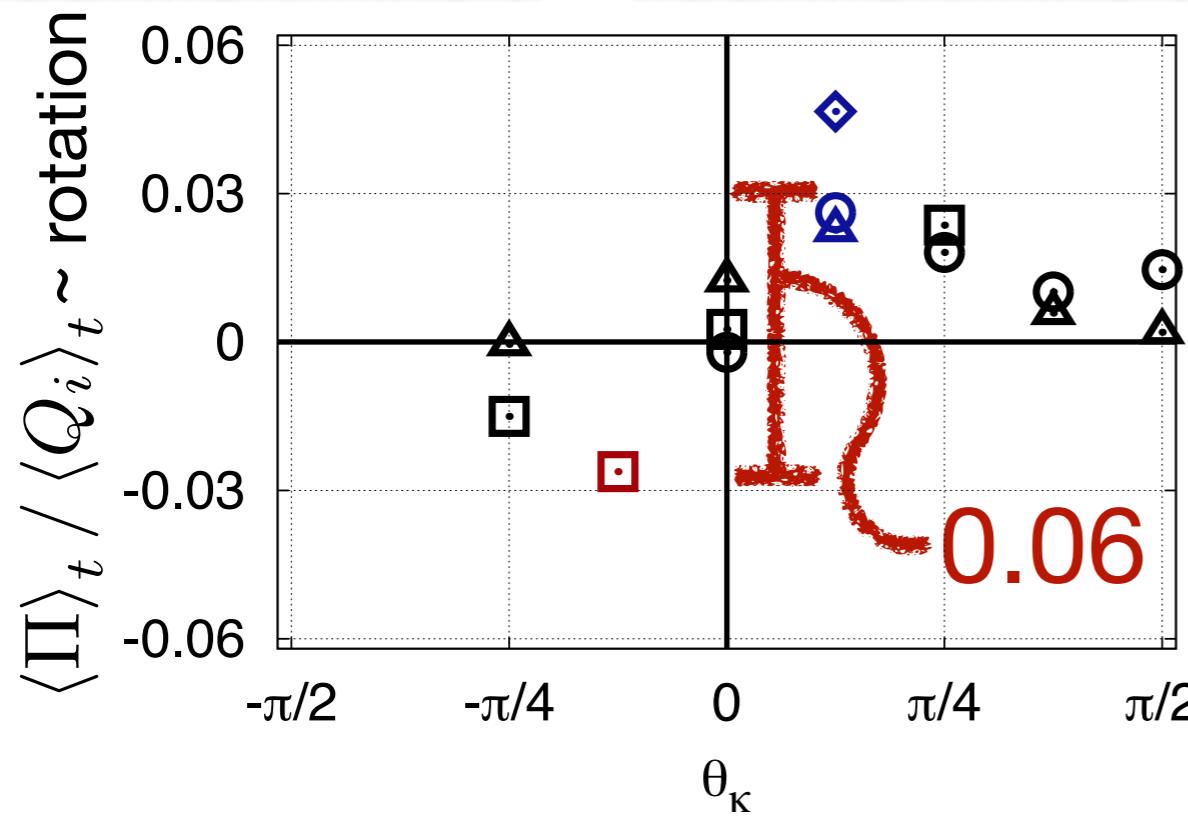
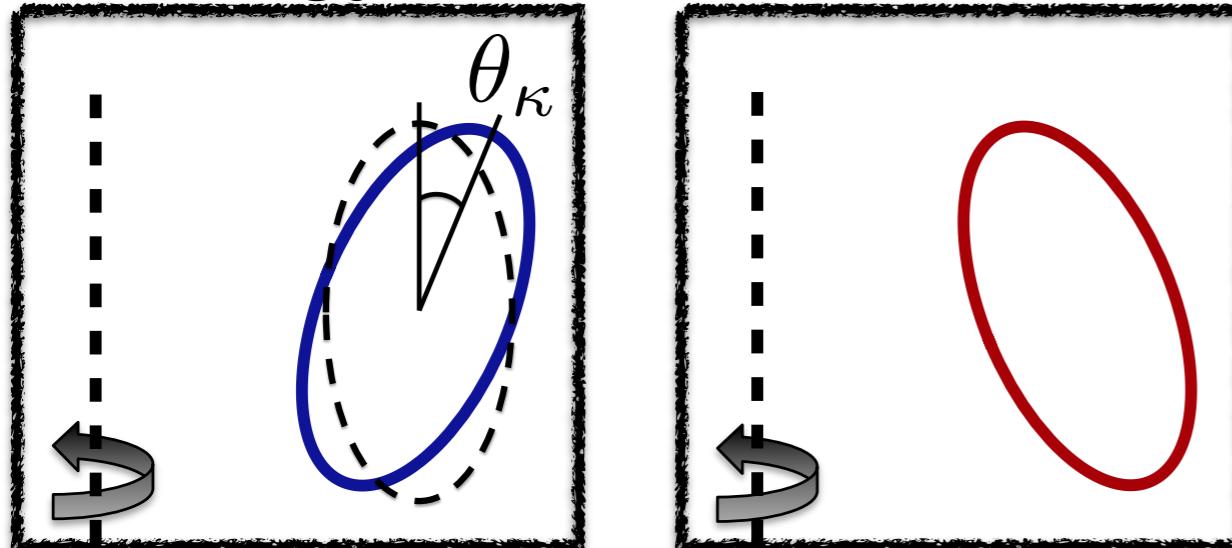


- Ignoring pinch is conservative, may enhance rotation by a factor of 3

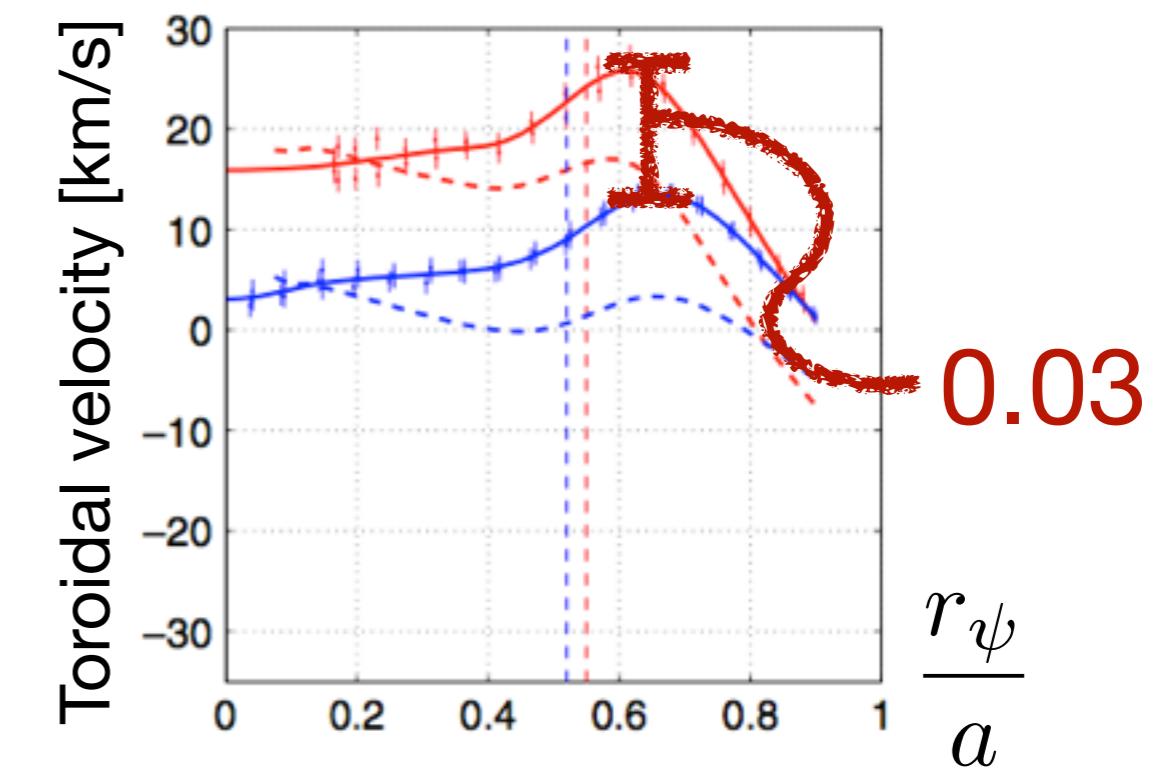
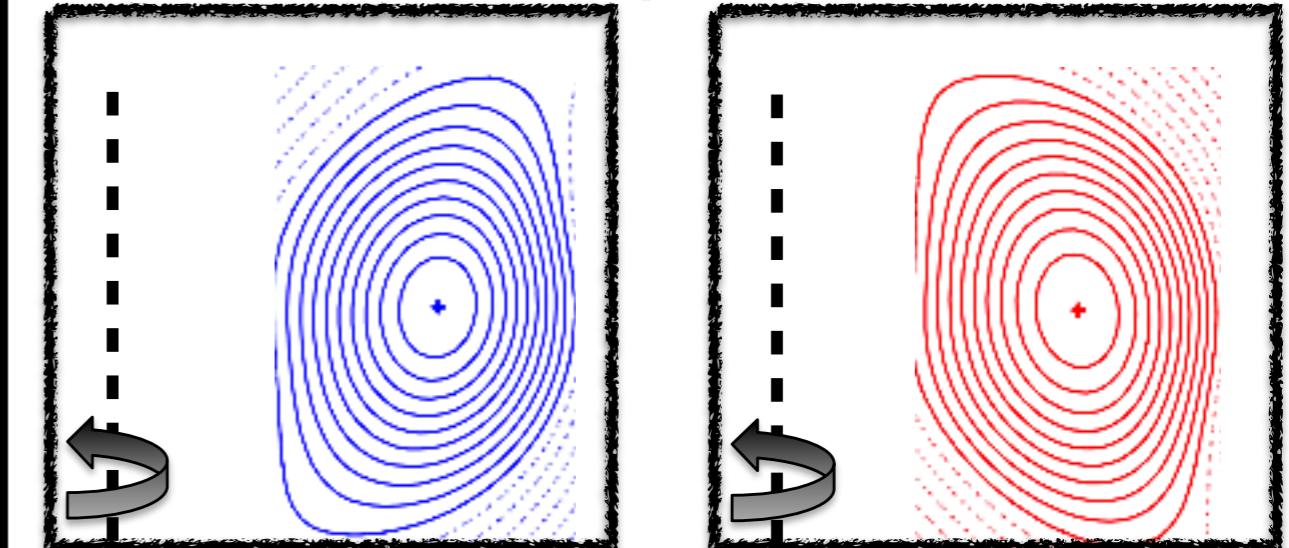
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# Outline

## Tokamaks

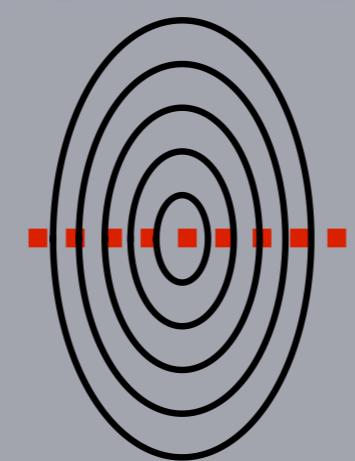
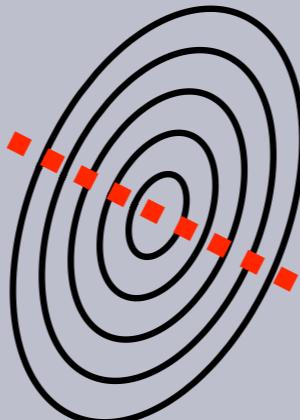
Mirror  
symmetric

$$M_A \approx 1.5\%$$

Up-down  
symmetric

$$M_A = 0$$

$$\langle \Pi(0, 0) \rangle_t = 0$$

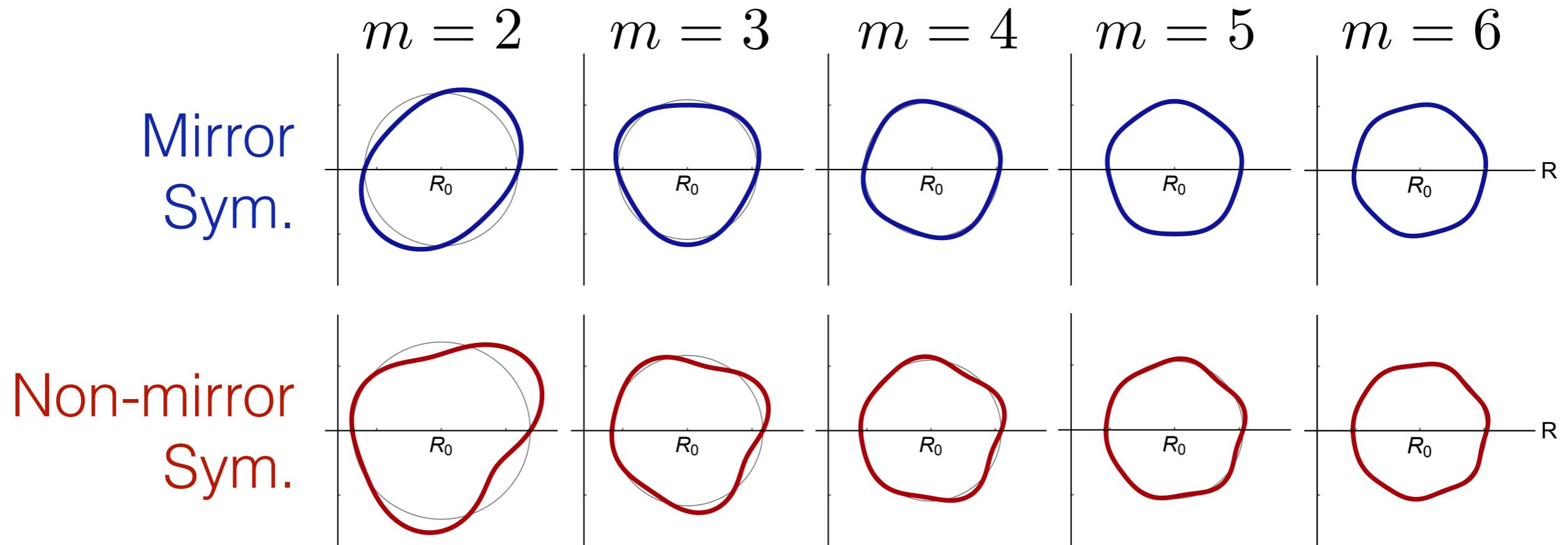


Non-mirror  
symmetric



# Scanning and expanding in $m \gg 1$

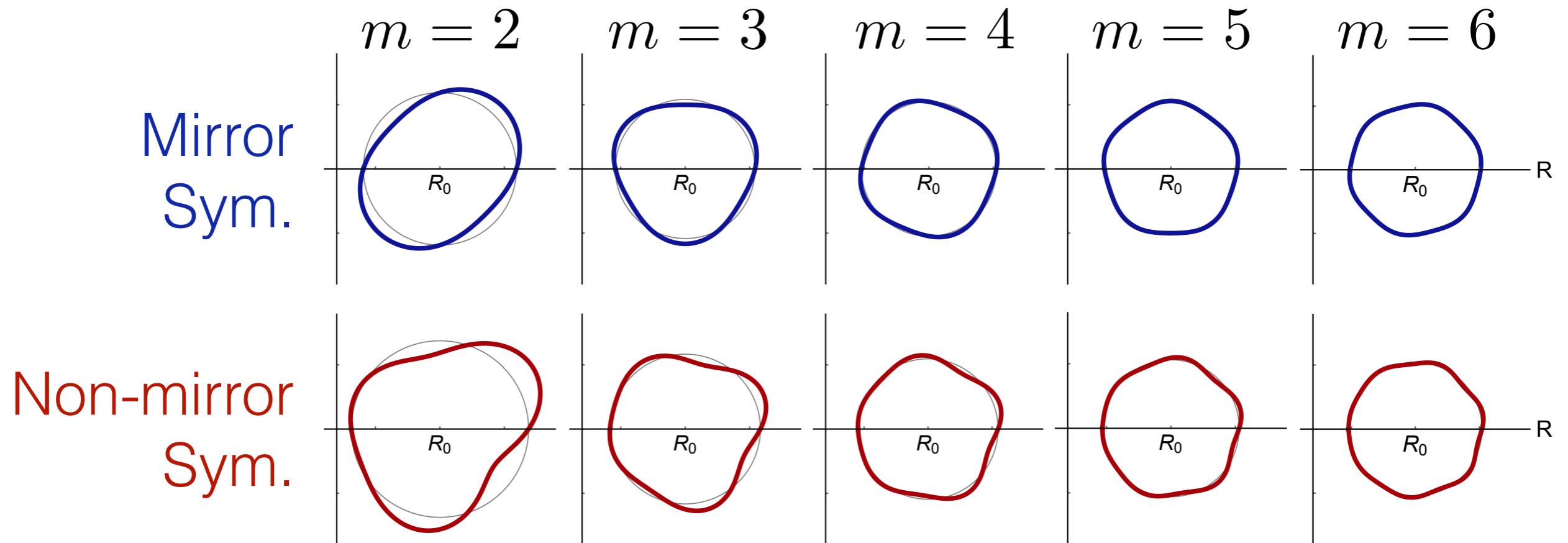
Ball, et al. *PPCF* (in prep).



- Counterintuitive because we just said low  $m$  penetrates better
- High  $m$  flux surfaces could transport momentum better
- Theoretical tool to allow analytic results

# Scaling of momentum flux with $m \gg 1$

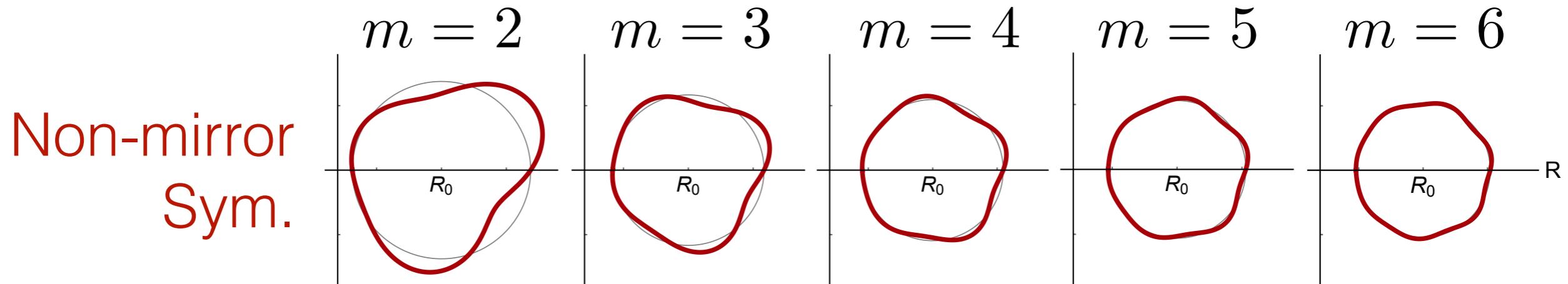
Ball, et al. PPCF (in prep).



- Expect that  $\langle \Pi(0, 0) \rangle_t \sim m^{-\alpha}$ , since the turbulent mode averages over the fast variation
- $\langle \Pi(0, 0) \rangle_t \sim m^{-\alpha}$  is true for non-mirror symmetric geometries, but not for mirror symmetric geometries

# Non-mirror symmetric geometries with $m \gg 1$

Ball, et al. *PPCF* (in prep).



- Created using two modes,  $m \sim n \gg 1$ , with distinct tilts  $\theta_{tm}$ ,  $\theta_{tn}$
- Assume reasonable shaping:  $C_m \sim C_n \sim m^{-2} \ll 1$
- Calculate geometric coefficients order-by-order in  $m \gg 1$
- Look for beating between fast shaping effects (creates symmetry-breaking on the connection length)

# Non-mirror symmetric geometries with $m \gg 1$

Ball, et al. *PPCF* (in prep).

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- Drift coefficient,  $v_{ds\alpha}$ , dominates (symmetry-breaking enters to lowest order in aspect ratio and is not a finite gyroradius effect)

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Ball, et al. *PPCF* (in prep).

- Drift coefficient,  $v_{ds\alpha}$ , dominates (symmetry-breaking enters to lowest order in aspect ratio and is not a finite gyroradius effect)

$$\begin{aligned} v_{ds\alpha} = & \frac{B_0}{R_0\Omega_s} \frac{dr_\psi}{d\psi} (\cos(\theta) + \hat{s}\theta\sin(\theta)) \\ & + \frac{B_0}{2R_0\Omega_s} \frac{dr_\psi}{d\psi} [(m^3 C_m^2 + n^3 C_n^2) \theta\sin(\theta) \\ & - \hat{s}\theta\cos(\theta) (mC_m\sin(m(\theta - \theta_{tm})) + nC_n\sin(n(\theta - \theta_{tn}))) \\ & + \hat{s}\theta\sin(\theta) (mC_m\cos(m(\theta - \theta_{tm})) + nC_n\cos(n(\theta - \theta_{tn}))) \\ & + mn \frac{m+n}{n-m} C_m C_n \sin(\theta) \\ & \times (\sin((n-m)\theta) \cos(m(\theta_{tm} - \theta_{tn})) - \cos((n-m)\theta) \sin(m(\theta_{tm} - \theta_{tn})))] \end{aligned}$$

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$$v_{ds\alpha} = \boxed{\frac{B_0}{R_0\Omega_s} \frac{dr_\psi}{d\psi} (\cos(\theta) + \hat{s}\theta\sin(\theta))} \quad \text{up-down sym.}$$
$$\boxed{+ \frac{B_0}{2R_0\Omega_s} \frac{dr_\psi}{d\psi} [(m^3 C_m^2 + n^3 C_n^2) \theta\sin(\theta)}$$
$$\boxed{- \hat{s}\theta\cos(\theta) (mC_m\sin(m(\theta - \theta_{tm})) + nC_n\sin(n(\theta - \theta_{tn})))}$$
$$\boxed{+ \hat{s}\theta\sin(\theta) (mC_m\cos(m(\theta - \theta_{tm})) + nC_n\cos(n(\theta - \theta_{tn})))}$$
$$\boxed{+ mn \frac{m+n}{n-m} C_m C_n \sin(\theta)} \quad \text{breaks symmetry}$$
$$\boxed{\times (\sin((n-m)\theta) \cos(m(\theta_{tm} - \theta_{tn})) - \cos((n-m)\theta) \sin(m(\theta_{tm} - \theta_{tn})))}]}$$

- Turbulence averages over high  $m$  variation

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$$+ \boxed{\frac{B_0}{2R_0\Omega_s} \frac{dr_\psi}{d\psi} [(m^3 C_m^2 + n^3 C_n^2) \theta\sin(\theta) - \hat{s}\theta\cos(\theta) (mC_m\sin(m(\theta - \theta_{tm})) + nC_n\sin(n(\theta - \theta_{tn}))) + \hat{s}\theta\sin(\theta) (mC_m\cos(m(\theta - \theta_{tm})) + nC_n\cos(n(\theta - \theta_{tn})))]} \quad \rightarrow 0$$
$$+ mn \frac{m+n}{n-m} C_m C_n \sin(\theta) \sim m^{-1} \times [(\sin((n-m)\theta) \cos(m(\theta_{tm} - \theta_{tn}))) - \cos((n-m)\theta) \sin(m(\theta_{tm} - \theta_{tn})))]$$

- Turbulence averages over high  $m$  variation

# Outline

## Tokamaks

Mirror  
symmetric

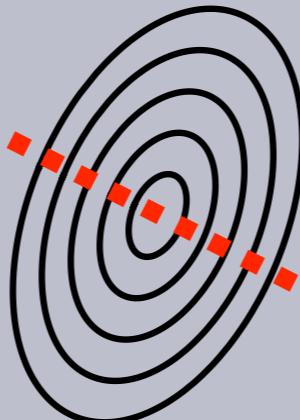
$$M_A \approx 1.5\%$$

$$\langle \Pi(0,0) \rangle_t \sim ???$$

Up-down  
symmetric

$$M_A = 0$$

$$\langle \Pi(0,0) \rangle_t = 0$$



Non-mirror  
symmetric



$$\langle \Pi(0,0) \rangle_t \sim m^{-1}$$

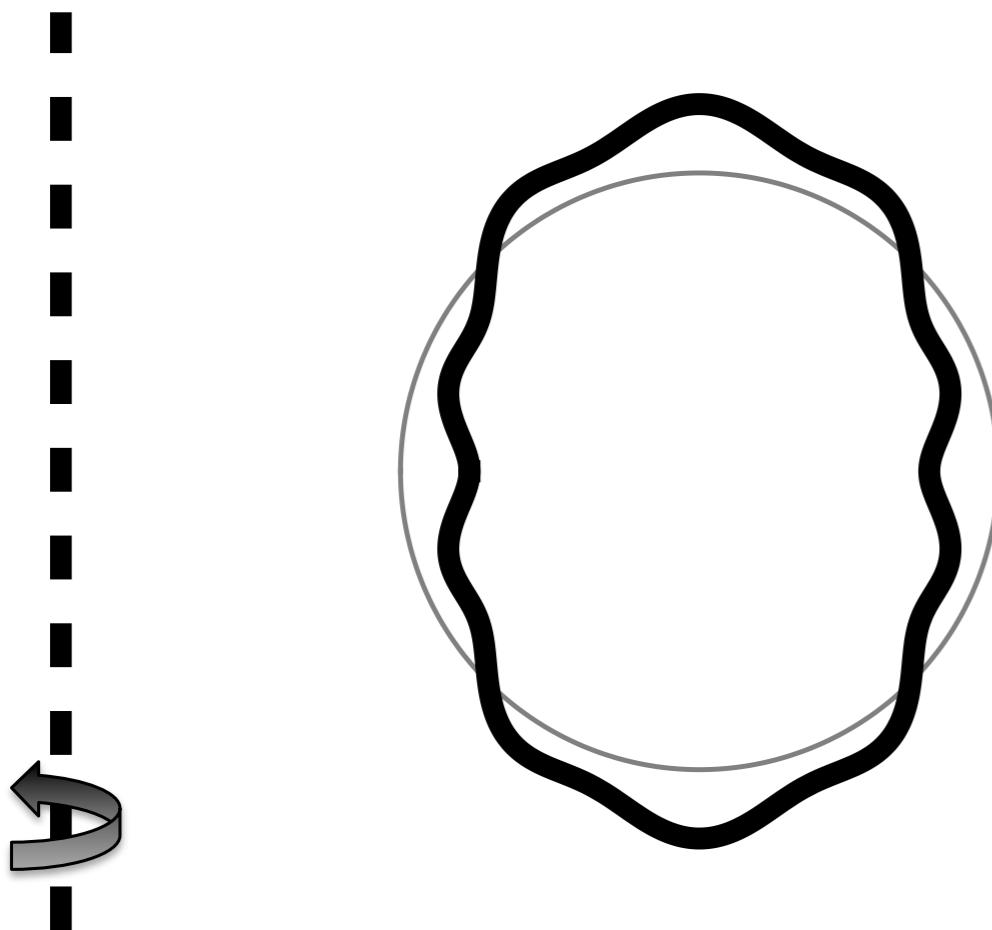
# Mirror symmetric geometries with $m \gg 1$

Ball, et al. *PPCF* (in prep).

- Specify geometry, separating a fast poloidal coordinate  $z \equiv m\theta$  from  $\theta$

## Up-down symmetric

e.g.  $r_0^{\text{ud}}(\theta, z) = r_{\psi 0} (1 - C_2 \cos(2\theta) - C_m \cos(m\theta))$



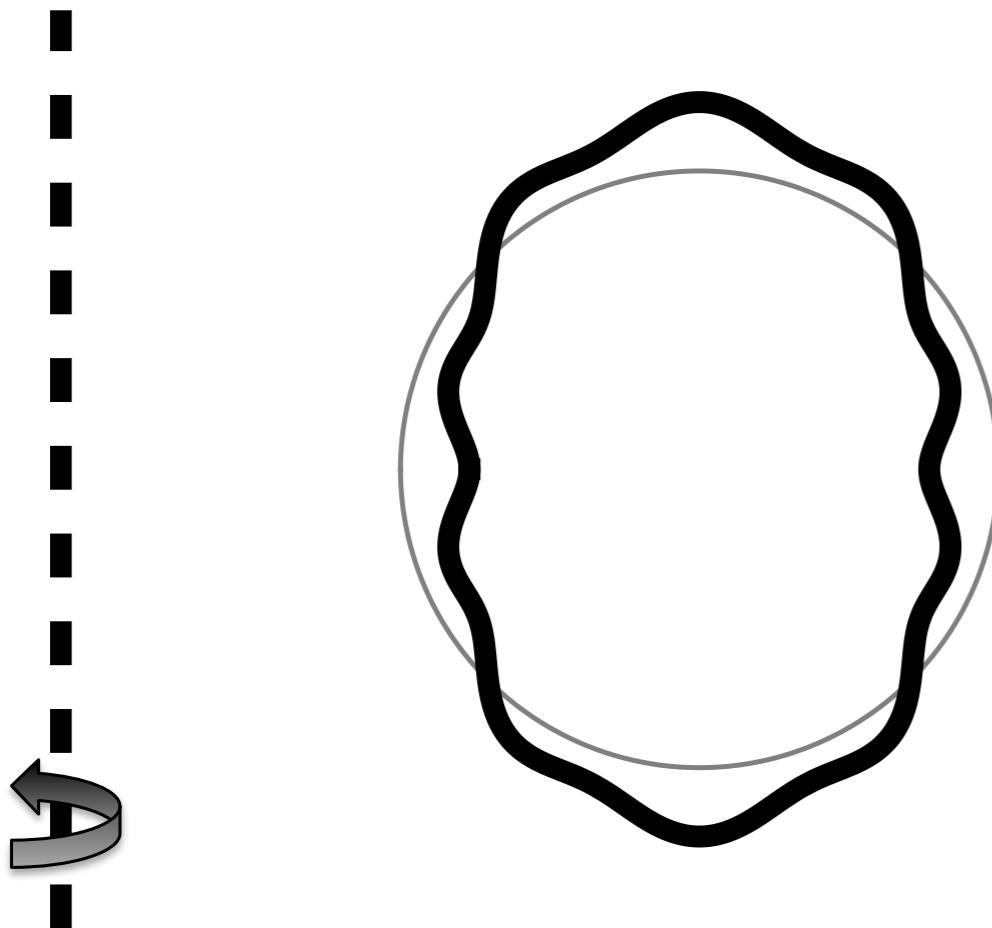
# Mirror symmetric geometries with $m \gg 1$

Ball, et al. *PPCF* (in prep).

- Create a second geometry by tilting the fast flux surface shaping

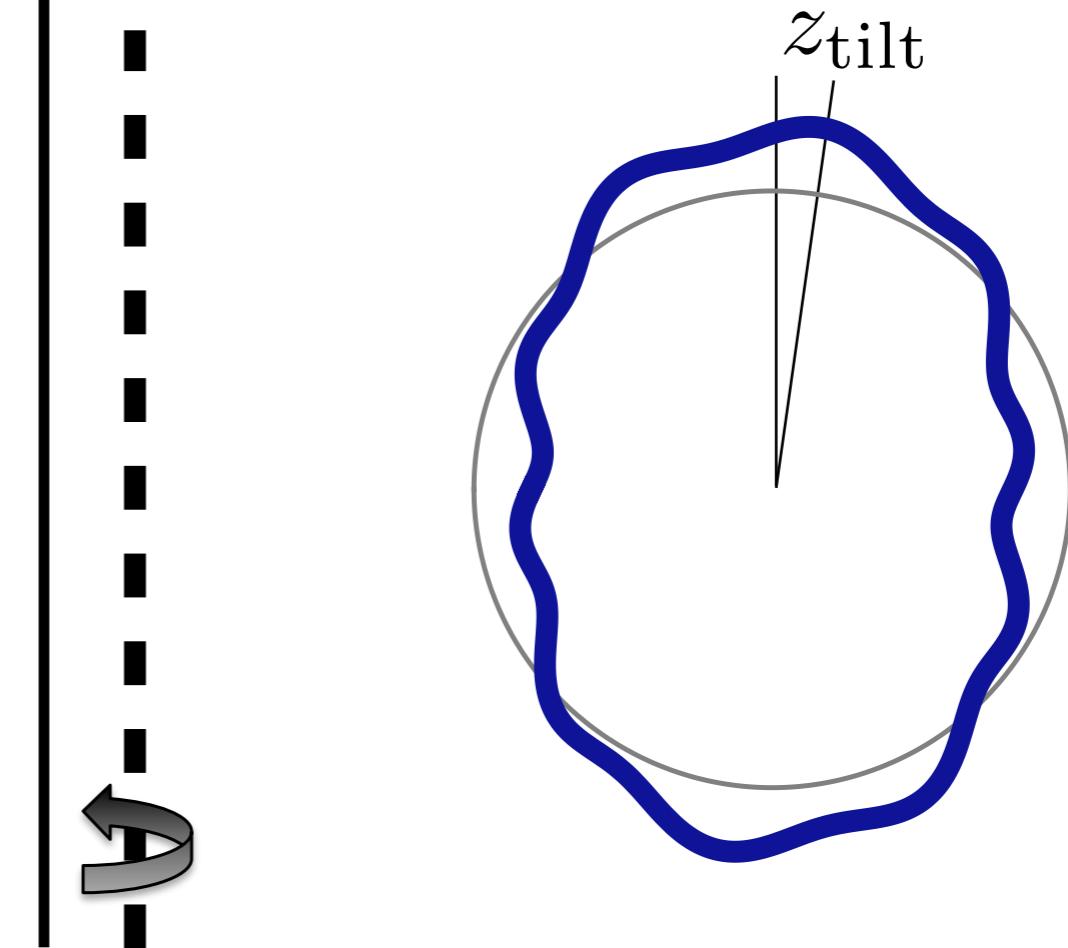
## Up-down symmetric

$$r_0^{\text{ud}}(\theta, z)$$



## Mirror symmetric

$$r_0^{\text{mir}}(\theta, z) = r_0^{\text{ud}}(\theta, z + z_{\text{tilt}})$$



# Mirror symmetric geometries with $m \gg 1$

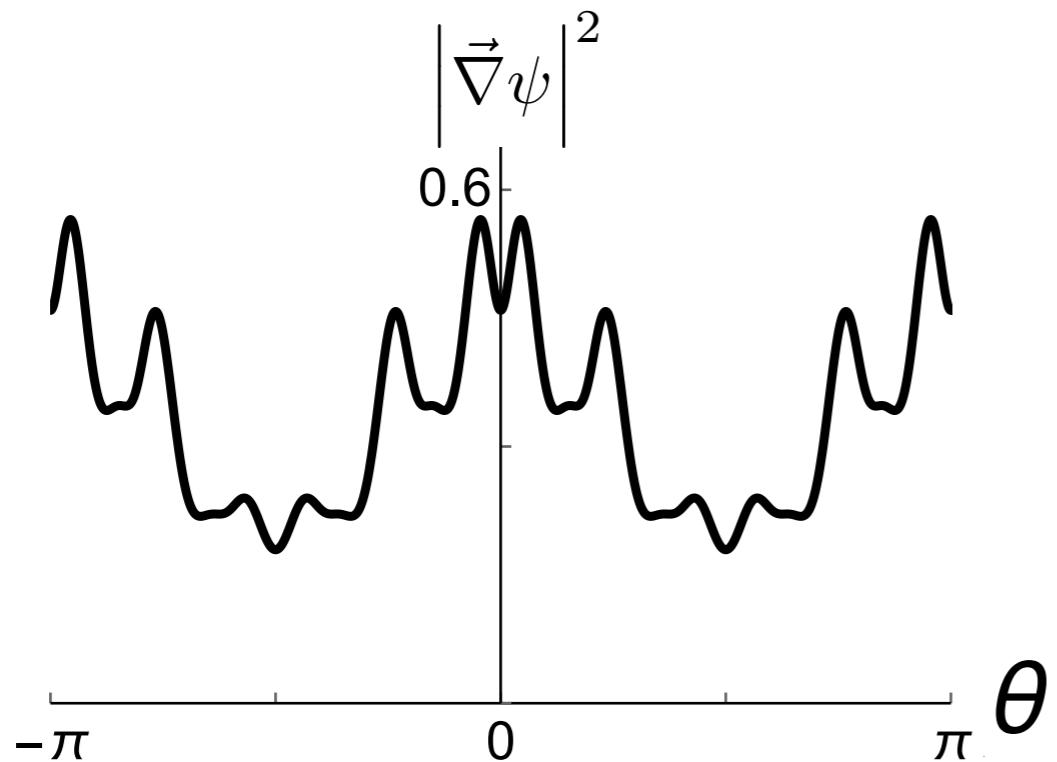
Ball, et al. *PPCF* (in prep).

- Look for differences in the geometric coefficients (critical and non-trivial)

## Up-down symmetric

$$r_0^{\text{ud}}(\theta, z)$$

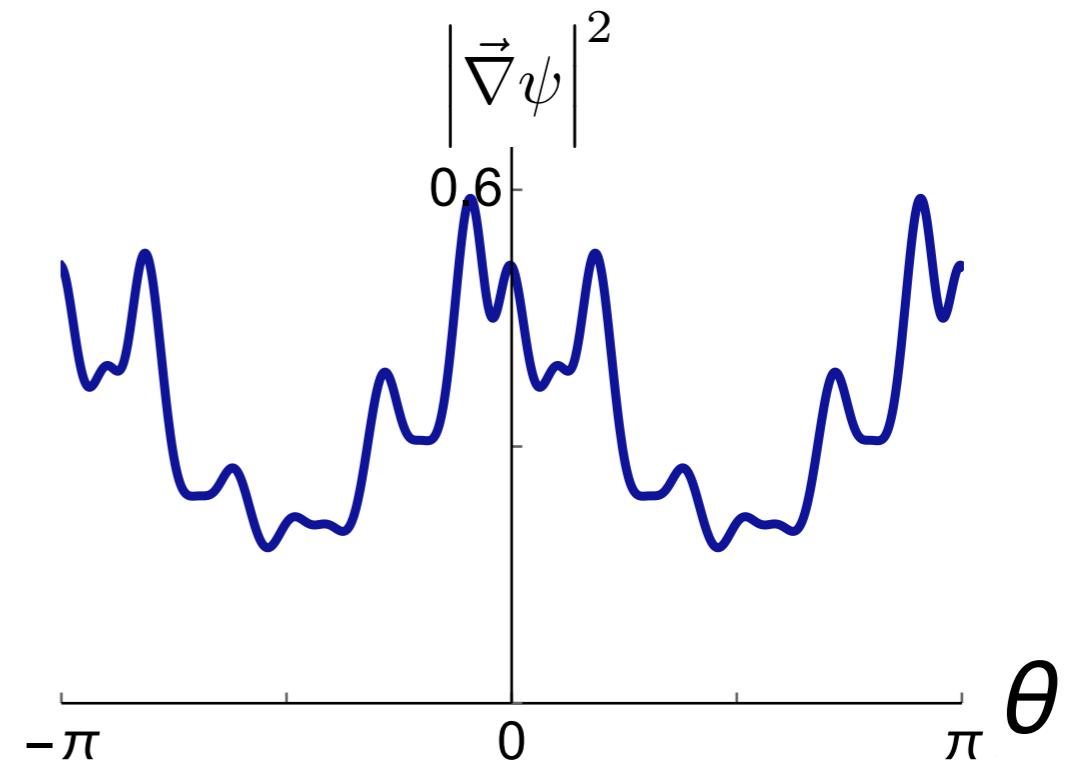
$$Q_{\text{geo}}^{\text{ud}}(\theta, z)$$



## Mirror symmetric

$$r_0^{\text{mir}}(\theta, z) = r_0^{\text{ud}}(\theta, z + z_{\text{tilt}})$$

$$Q_{\text{geo}}^{\text{mir}}(\theta, z) = Q_{\text{geo}}^{\text{ud}}(\theta, z + z_{\text{tilt}})$$



# Mirror symmetric geometries with $m \gg 1$

Ball, et al. *PPCF* (in prep).

- The geometric coefficients are the only way the geometry enters, so...

## Up-down symmetric

$$r_0^{\text{ud}}(\theta, z)$$

$$Q_{\text{geo}}^{\text{ud}}(\theta, z)$$

$$h_s^{\text{ud}}(\theta, z)$$

$$\phi^{\text{ud}}(\theta, z)$$

## Mirror symmetric

$$r_0^{\text{mir}}(\theta, z) = r_0^{\text{ud}}(\theta, z + z_{\text{tilt}})$$

$$Q_{\text{geo}}^{\text{mir}}(\theta, z) = Q_{\text{geo}}^{\text{ud}}(\theta, z + z_{\text{tilt}})$$

$$h_s^{\text{mir}}(\theta, z) = h_s^{\text{ud}}(\theta, z + z_{\text{tilt}})$$

$$\phi^{\text{mir}}(\theta, z) = \phi^{\text{ud}}(\theta, z + z_{\text{tilt}})$$

# Mirror symmetric geometries with $m \gg 1$

Ball, et al. *PPCF* (in prep).

- Define  $\pi_\zeta$  such that  $\Pi = \langle\langle \pi_\zeta(\theta, z) \rangle_z \rangle_\psi$

## Up-down symmetric

$$r_0^{\text{ud}}(\theta, z)$$

$$Q_{\text{geo}}^{\text{ud}}(\theta, z)$$

$$h_s^{\text{ud}}(\theta, z)$$

$$\phi^{\text{ud}}(\theta, z)$$

$$\pi_\zeta^{\text{ud}}(\theta, z)$$

## Mirror symmetric

$$r_0^{\text{mir}}(\theta, z) = r_0^{\text{ud}}(\theta, z + z_{\text{tilt}})$$

$$Q_{\text{geo}}^{\text{mir}}(\theta, z) = Q_{\text{geo}}^{\text{ud}}(\theta, z + z_{\text{tilt}})$$

$$h_s^{\text{mir}}(\theta, z) = h_s^{\text{ud}}(\theta, z + z_{\text{tilt}})$$

$$\phi^{\text{mir}}(\theta, z) = \phi^{\text{ud}}(\theta, z + z_{\text{tilt}})$$

$$\pi_\zeta^{\text{mir}}(\theta, z) = \pi_\zeta^{\text{ud}}(\theta, z + z_{\text{tilt}})$$

# Mirror symmetric geometries with $m \gg 1$

Ball, et al. *PPCF* (in prep).

- The  $z$ -average in  $\Pi = \langle\langle \pi_\zeta(\theta, z) \rangle_z \rangle_\psi$  doesn't care about the tilt

## Up-down symmetric

$$r_0^{\text{ud}}(\theta, z)$$

$$Q_{\text{geo}}^{\text{ud}}(\theta, z)$$

$$h_s^{\text{ud}}(\theta, z)$$

$$\phi^{\text{ud}}(\theta, z)$$

$$\pi_\zeta^{\text{ud}}(\theta, z)$$

$$\Pi^{\text{ud}} = 0$$

## Mirror symmetric

$$r_0^{\text{mir}}(\theta, z) = r_0^{\text{ud}}(\theta, z + z_{\text{tilt}})$$

$$Q_{\text{geo}}^{\text{mir}}(\theta, z) = Q_{\text{geo}}^{\text{ud}}(\theta, z + z_{\text{tilt}})$$

$$h_s^{\text{mir}}(\theta, z) = h_s^{\text{ud}}(\theta, z + z_{\text{tilt}})$$

$$\phi^{\text{mir}}(\theta, z) = \phi^{\text{ud}}(\theta, z + z_{\text{tilt}})$$

$$\pi_\zeta^{\text{mir}}(\theta, z) = \pi_\zeta^{\text{ud}}(\theta, z + z_{\text{tilt}})$$

$$\Pi^{\text{mir}} = 0?$$

# Mirror symmetric geometries with $m \gg 1$

Ball, et al. *PPCF* (in prep).

- We relied on expansion in  $m \gg 1$  to separate poloidal scales

## Up-down symmetric

$$r_0^{\text{ud}}(\theta, z)$$

$$Q_{\text{geo}}^{\text{ud}}(\theta, z)$$

$$h_s^{\text{ud}}(\theta, z)$$

$$\phi^{\text{ud}}(\theta, z)$$

$$\pi_\zeta^{\text{ud}}(\theta, z)$$

$$\Pi^{\text{ud}} = 0$$

## Mirror symmetric

$$r_0^{\text{mir}}(\theta, z) = r_0^{\text{ud}}(\theta, z + z_{\text{tilt}})$$

$$Q_{\text{geo}}^{\text{mir}}(\theta, z) = Q_{\text{geo}}^{\text{ud}}(\theta, z + z_{\text{tilt}})$$

$$h_s^{\text{mir}}(\theta, z) = h_s^{\text{ud}}(\theta, z + z_{\text{tilt}})$$

$$\phi^{\text{mir}}(\theta, z) = \phi^{\text{ud}}(\theta, z + z_{\text{tilt}})$$

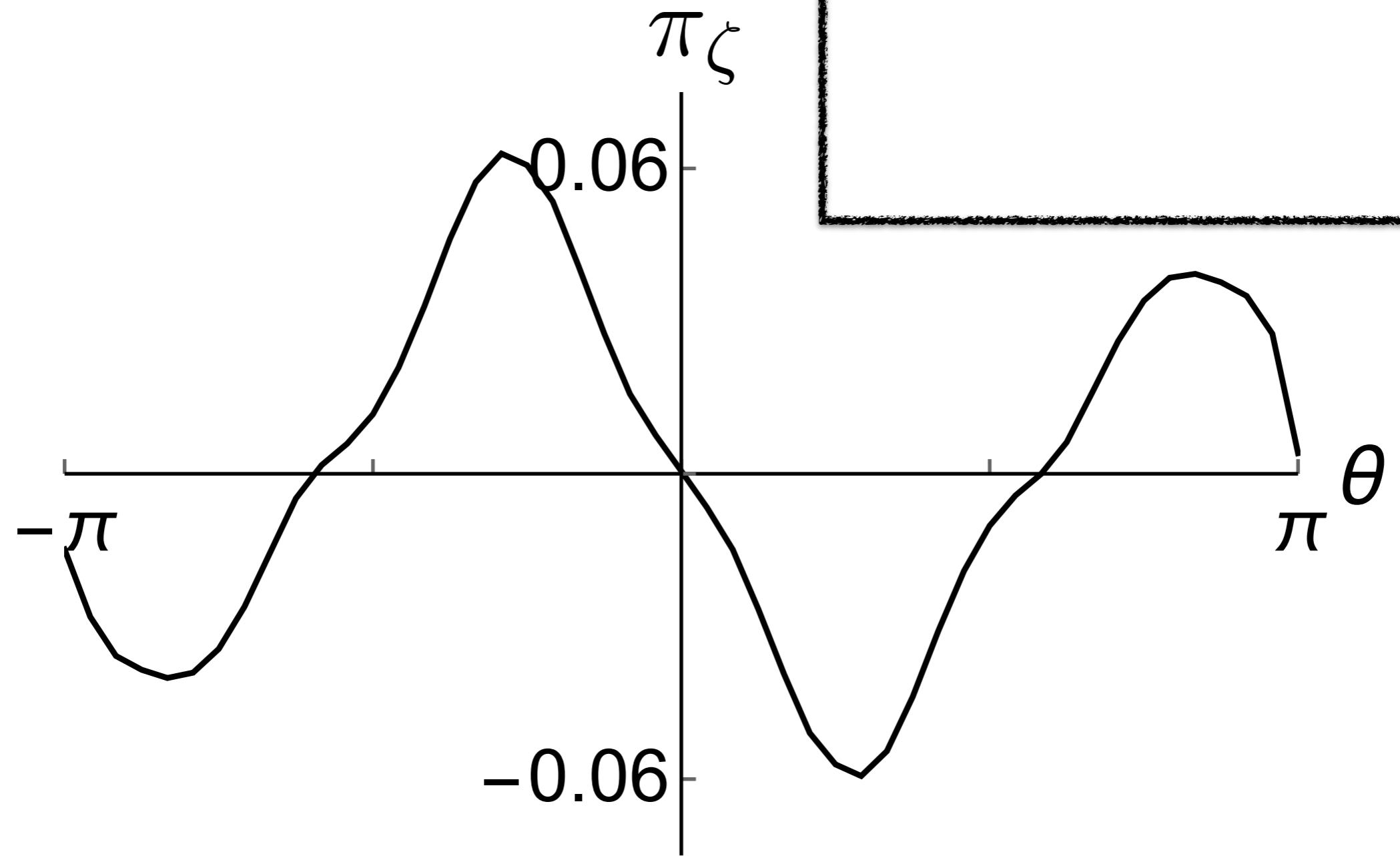
$$\pi_\zeta^{\text{mir}}(\theta, z) = \pi_\zeta^{\text{ud}}(\theta, z + z_{\text{tilt}})$$

~~$$\Pi^{\text{mir}} = 0?$$~~

$$\Pi \lesssim \exp(-\alpha m)$$

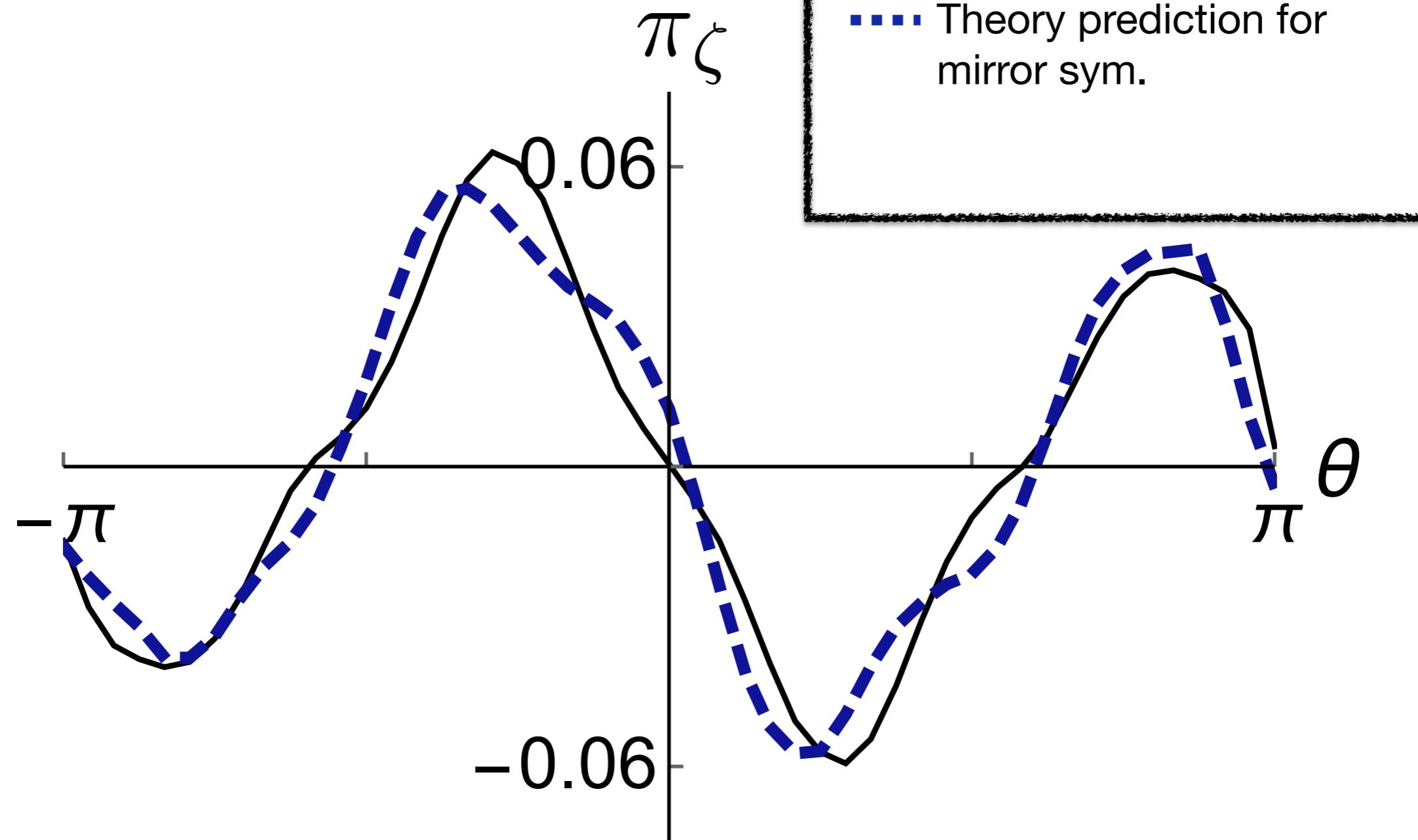
Mom. flux distribution, testing  $\pi_{\zeta}^{\text{mir}}(\theta, z) = \pi_{\zeta}^{\text{ud}}(\theta, z + z_{\text{tilt}})$   
Ball, et al. *PPCF* (in prep).

m=7 geometry:



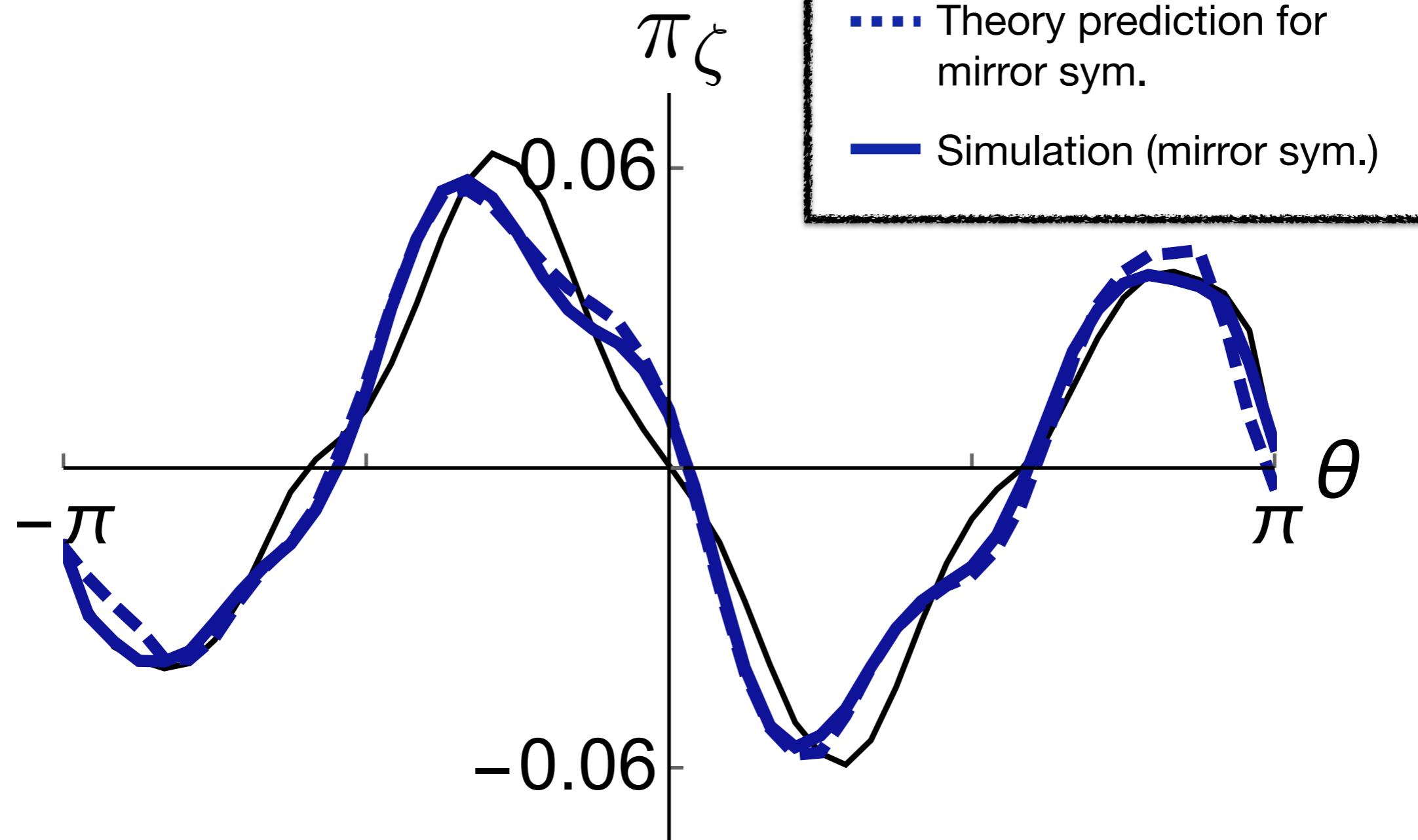
Mom. flux distribution, testing  $\pi_{\zeta}^{\text{mir}}(\theta, z) = \pi_{\zeta}^{\text{ud}}(\theta, z + z_{\text{tilt}})$   
Ball, et al. *PPCF* (in prep).

m=7 geometry:



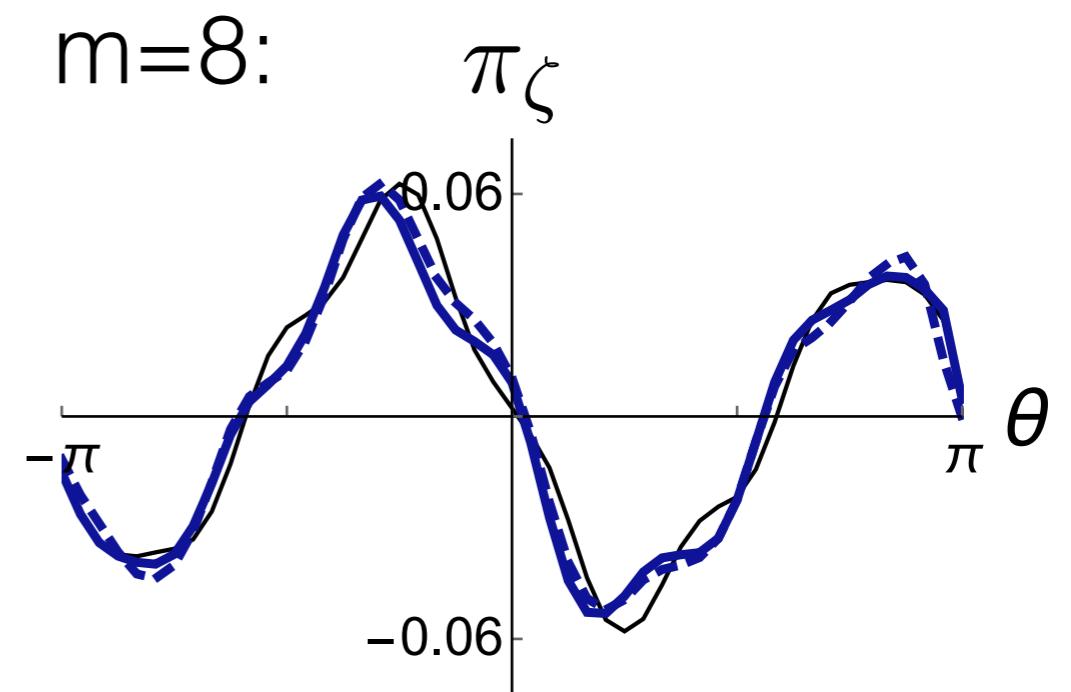
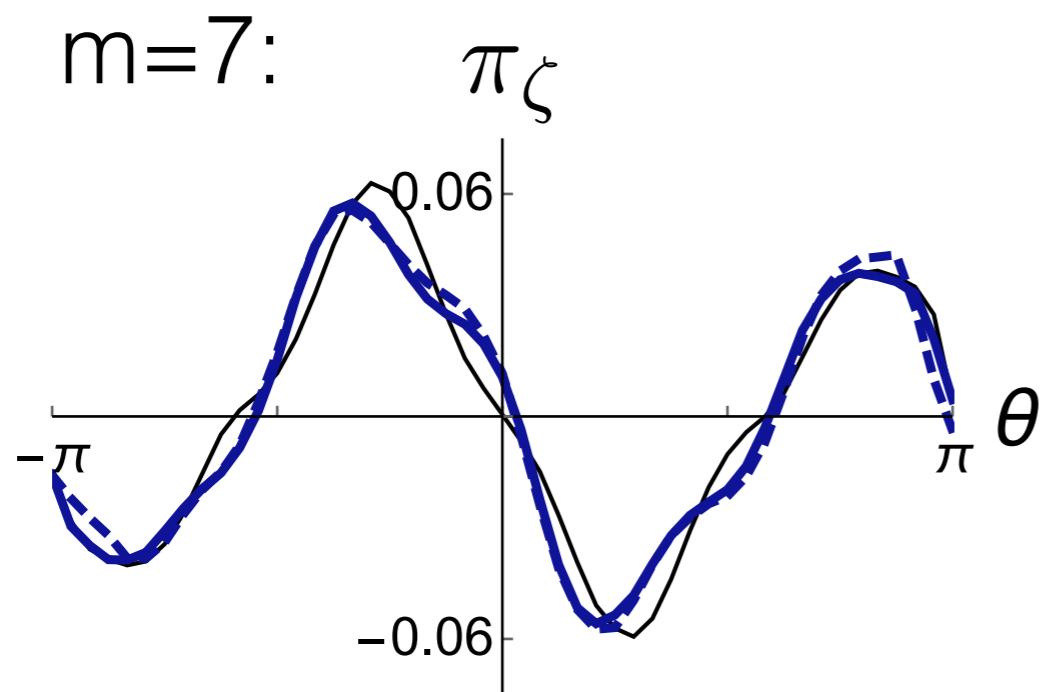
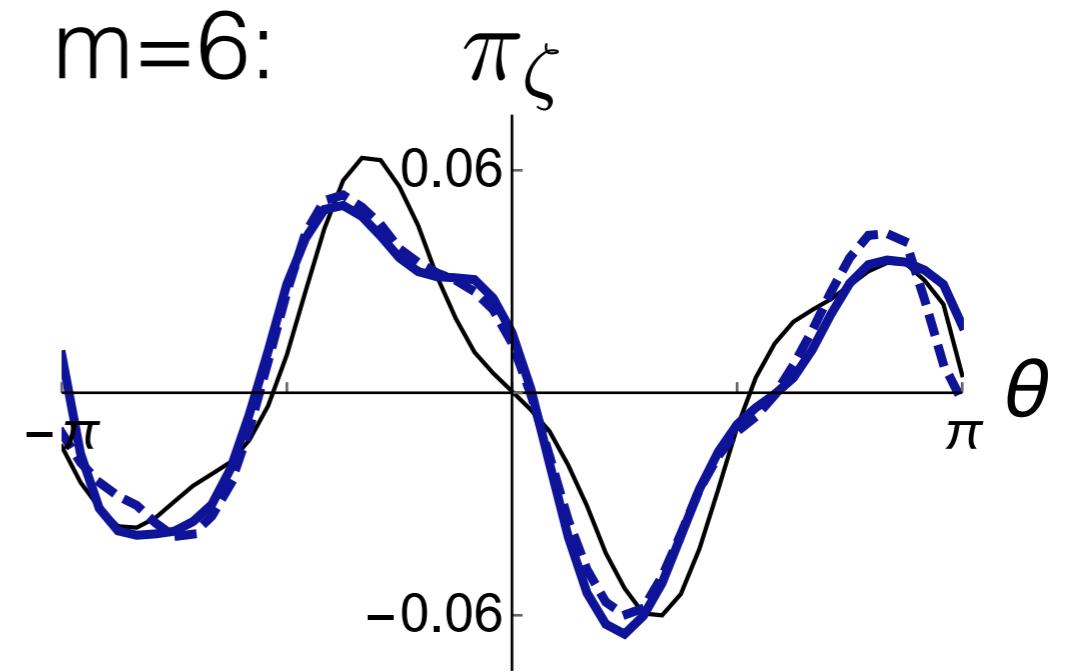
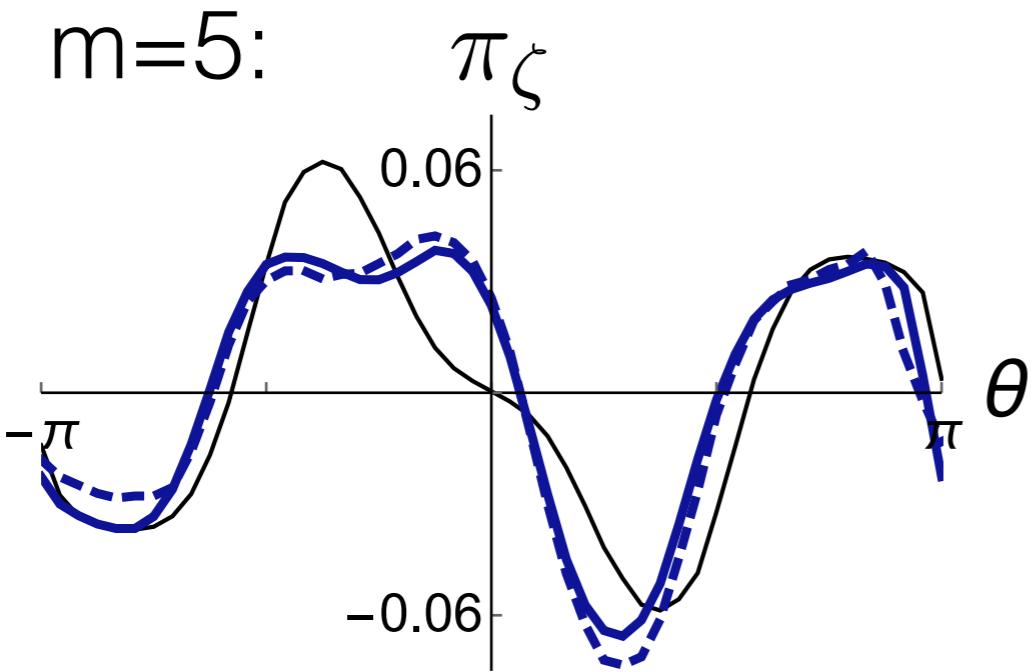
Mom. flux distribution, testing  $\pi_{\zeta}^{\text{mir}}(\theta, z) = \pi_{\zeta}^{\text{ud}}(\theta, z + z_{\text{tilt}})$   
Ball, et al. *PPCF* (in prep).

m=7 geometry:



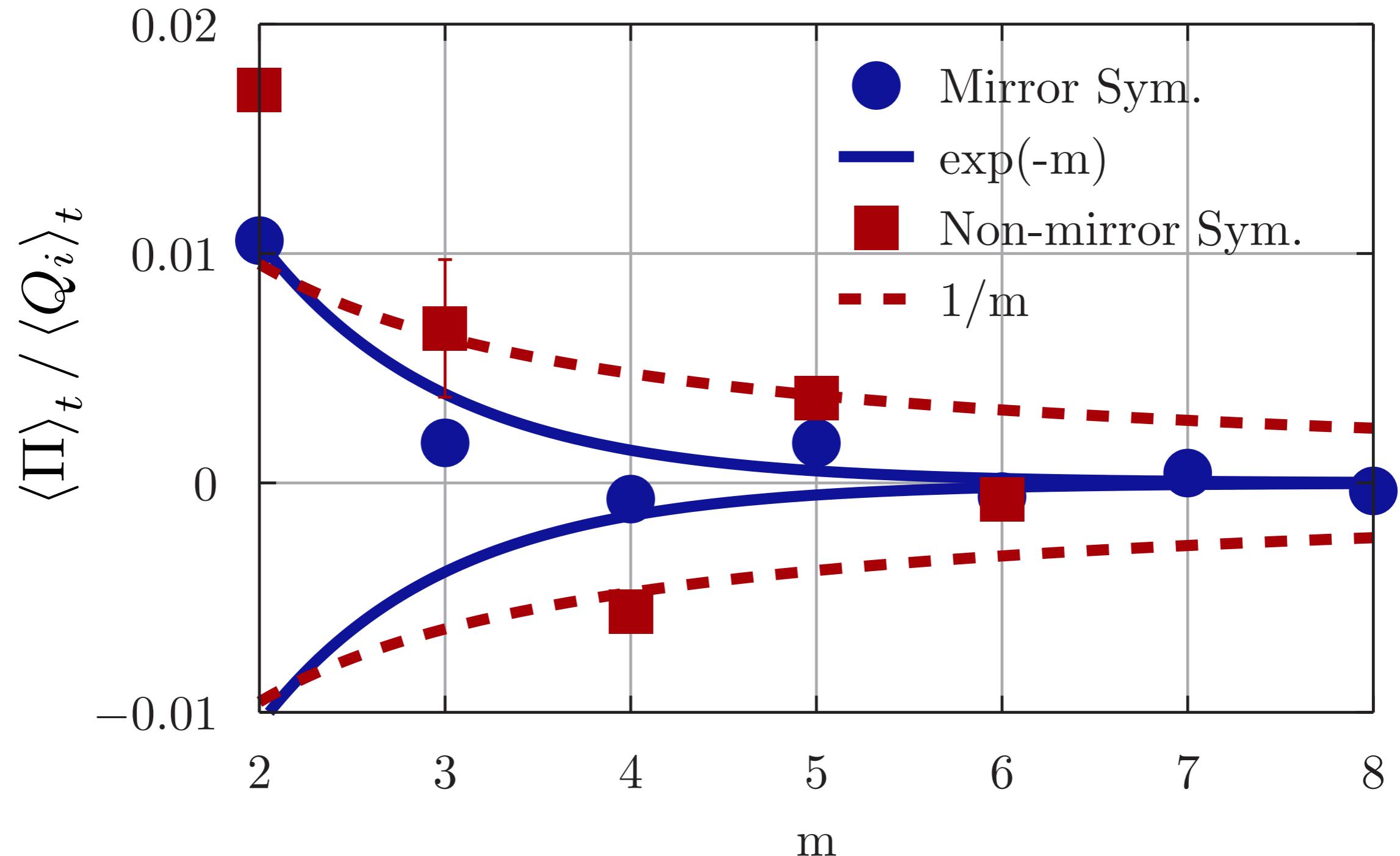
# Mom. flux distribution, testing $\pi_{\zeta}^{\text{mir}}(\theta, z) = \pi_{\zeta}^{\text{ud}}(\theta, z + z_{\text{tilt}})$

Ball, et al. *PPCF* (in prep).



# Numerical scaling with $m \gg 1$

Ball, et al. PPCF (in prep).



# Outline

## Tokamaks

Mirror  
symmetric

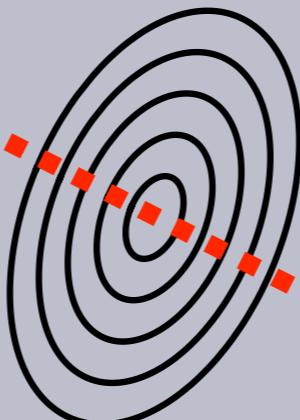
$$M_A \approx 1.5\%$$

$$\langle \Pi(0,0) \rangle_t \lesssim \exp(-\alpha m)$$

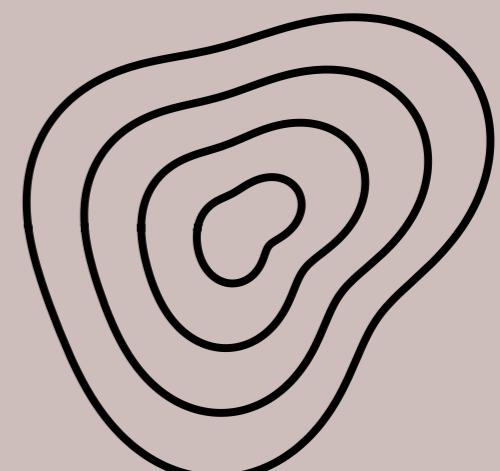
Up-down  
symmetric

$$M_A = 0$$

$$\langle \Pi(0,0) \rangle_t = 0$$



Non-mirror  
symmetric

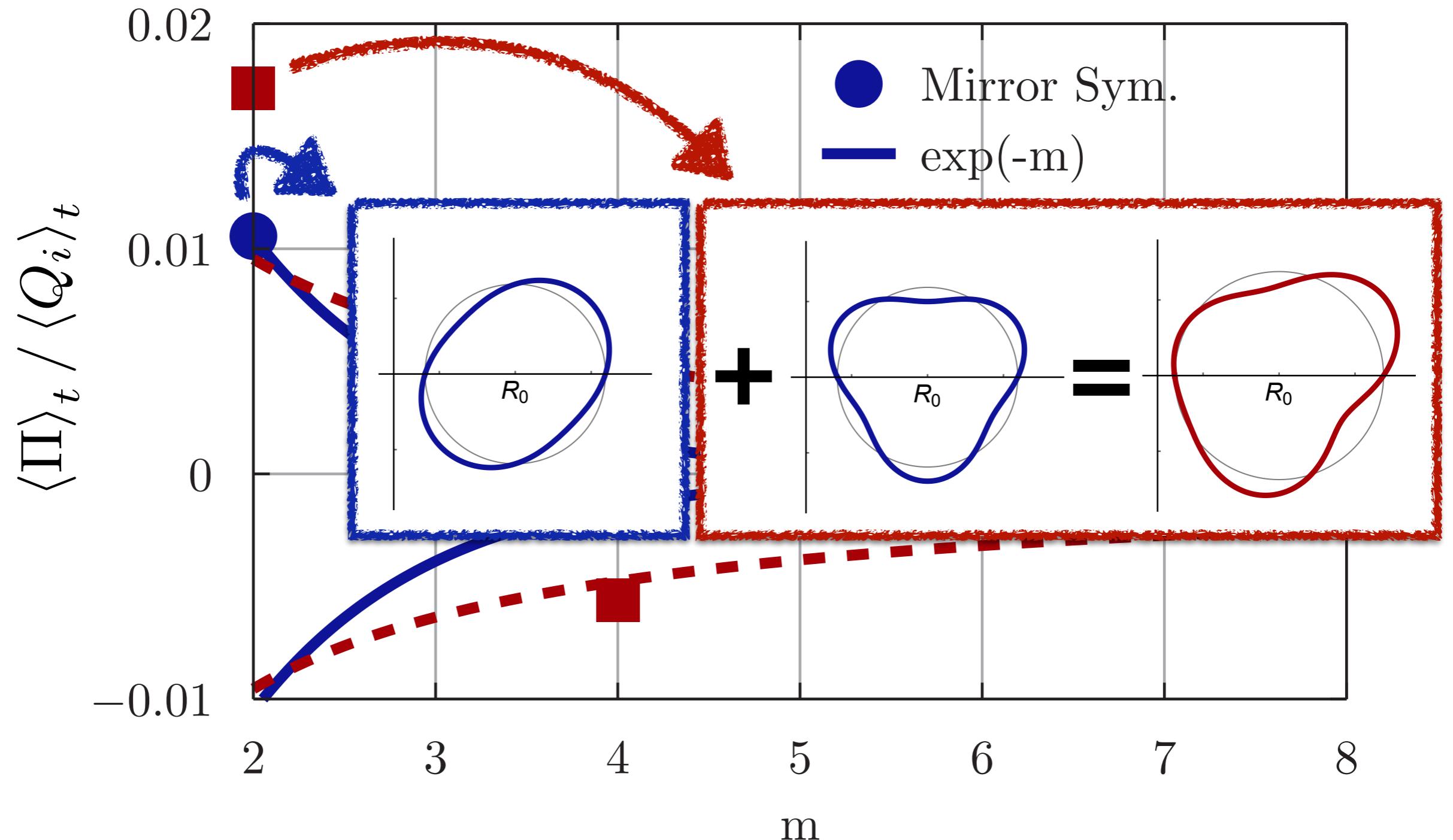


$$\langle \Pi(0,0) \rangle_t \sim m^{-1}$$

$M_A \approx ???$

# Numerical scaling with $m \gg 1$

Ball, et al. PPCF (in prep).



# Outline

## Tokamaks

Mirror  
symmetric

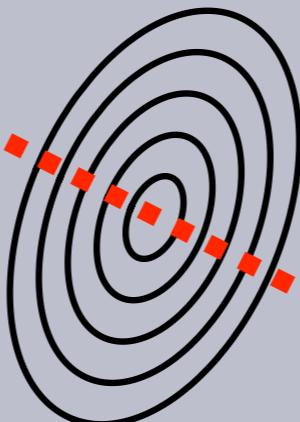
$$M_A \approx 1.5\%$$

$$\langle \Pi(0,0) \rangle_t \lesssim \exp(-\alpha m)$$

Up-down  
symmetric

$$M_A = 0$$

$$\langle \Pi(0,0) \rangle_t = 0$$



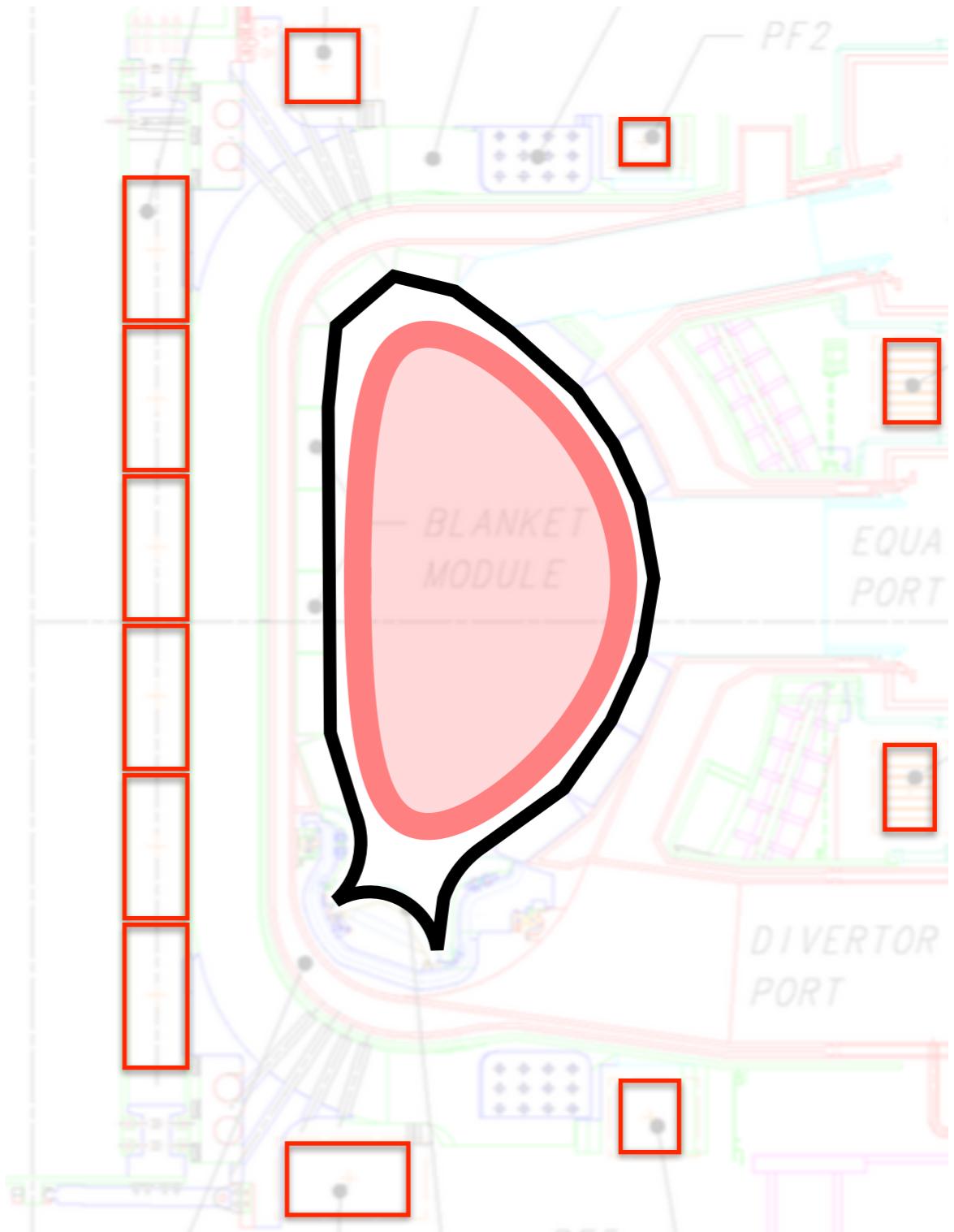
Non-mirror  
symmetric



$$\langle \Pi(0,0) \rangle_t \sim m^{-1}$$
$$M_A \approx 2.5\%?$$

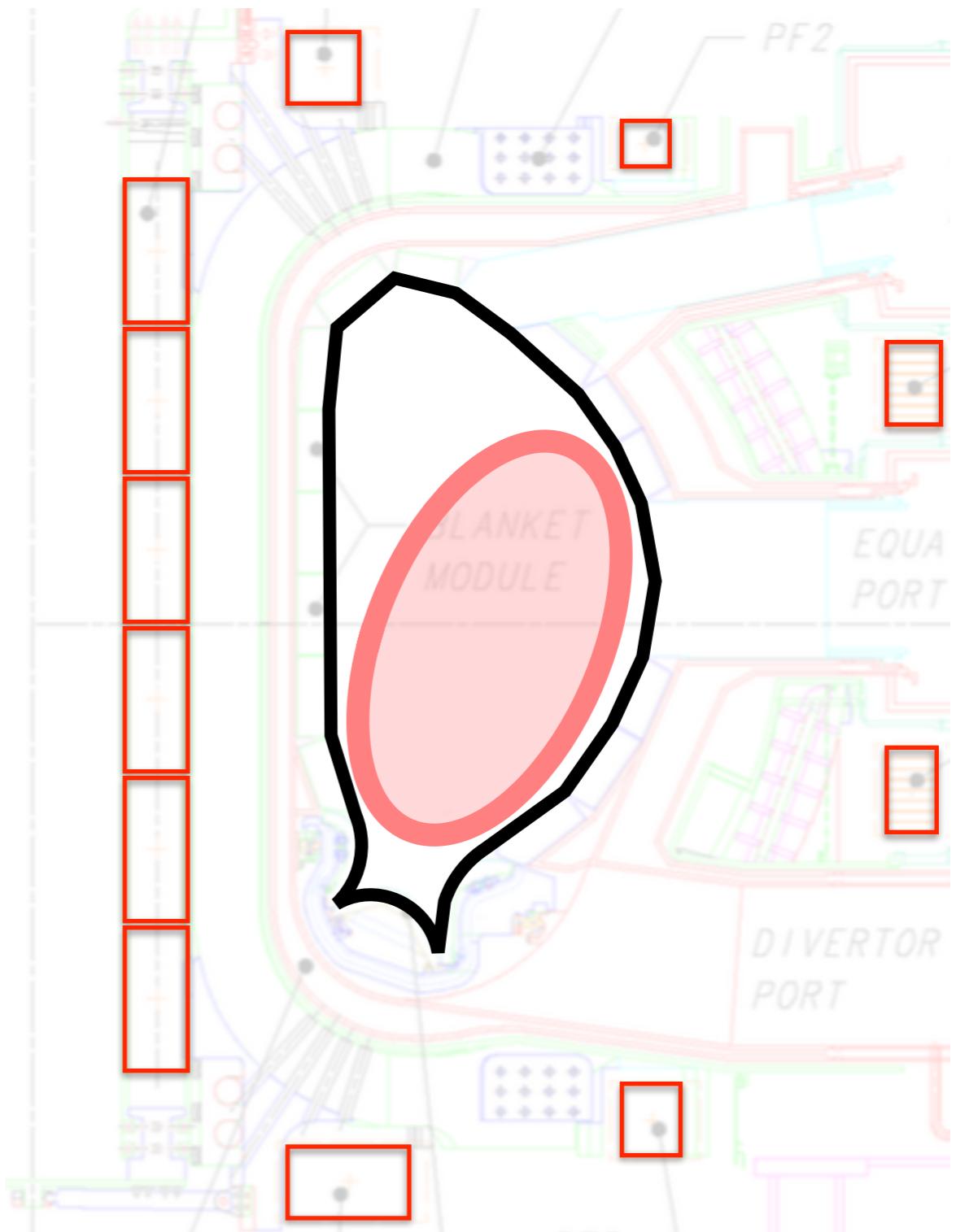
# Can something like this be done in ITER?

- I think so, but there is a catch
- The shape of the first wall is fixed
- Each shaping coils has a current limit



# Can something like this be done in ITER?

- I think so, but there is a catch
  - The shape of the first wall is fixed
- ➡ Reduced plasma volume
- Each shaping coils has a current limit
- ➡ Reduced plasma current
- For  $\beta_N = 3$ , it's worth a shot



# Conclusions

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- Intrinsic rotation generated by up-down asymmetry scales well to larger machines (ITER, DEMO, etc.), unlike other mechanisms
- Tilting the elongation of flux surfaces by  $\sim 20^\circ$  is a simple way to generate significant rotation
  - The magnitude of rotation is roughly what is needed to allow  $\beta_N = 3$  in ITER
- Breaking mirror symmetry, in addition to up-down symmetry, has the potential to create even more rotation

Thank you!

# Summary

## Tokamaks

Mirror  
symmetric

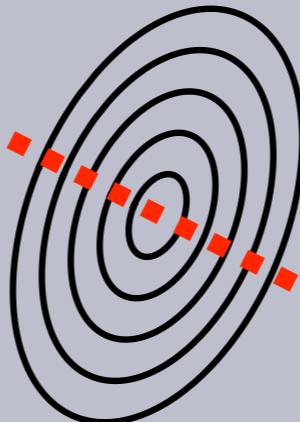
$$M_A \approx 1.5\%$$

$$\langle \Pi(0,0) \rangle_t \lesssim \exp(-\alpha m)$$

Up-down  
symmetric

$$M_A = 0$$

$$\langle \Pi(0,0) \rangle_t = 0$$



Non-mirror  
symmetric



$$\langle \Pi(0,0) \rangle_t \sim m^{-1}$$
$$M_A \approx 2.5\%?$$